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### Oak Ridge National Laboratory

Some (selective) transfer opportunities at FRIB

• Transfer reactions for neutron capture

S.D. Pain

- Constraining n-capture/Surrogate reactions with fast beams
- Transfer reactions on isomers
  - sd-shell nuclides
  - Constraining proton capture [(d,p) for (p,γ)]
  - Direct ( $\alpha$ ,p)
  - Comparison with SM





### FRIB TA Workshop, May 2023

## Transfer reaction observables

- Level energies (few keV hundreds keV, depending on reaction, beam and instrumentation)
- (Differential) cross sections
  - transferred orbital angular momenta (parity, some J)
- Decay channels of excited states
  - γ spec
  - $\gamma$ /particle emission probabilities [discrete, or P(E<sub>x</sub>)]
- Direct reaction analysis
  - Spectroscopic factors (C<sup>2</sup>S)
    - Structure
    - Astrophysics
- Surrogate reactions



Selective of particle/hole states

- Stripping reactions [eg (d,p) (d,n)] probe of single-particle excitations
- Pickup reactions [eg (p,d) (d,<sup>3</sup>He)] probe of hole excitations
- Spin through momentum matching (choice of probe, and beam energy)

### Astrophysics

Level energies (exp. affect rates)

Spins (barrier penetrabilities – ell, not J, most important)

C<sup>2</sup>S

Constraints on LD/GSF – eg SRM

Preferentially populates particle excitations (depending on WF, may or may not see via other probes – beta decay, KO, etc)

#### Branching ratios

Though direct measurements are goal, need to discover important states first by some other means

## Transfer reactions at FRIB

### Beam energy ~3-50 MeV/A

- Cross sections (absolute and differential)
- Beam intensity
- Kinematic compression
- Beam optics
- Special cases...



150

10

4 MeV/A

6 MeV/A

—10 MeV/A

20 MeV/A

30 MeV/A

40 MeV/A

15



Detectors at FRIB

Si+Solenoid arrays (SOLARIS,



Active targets (ATTPC)



Gas jet targets (*JENSA*)

Devices often coupled to recoil separators (S800, SECAR, ISLA)



Charged-particle arrays (*HiRA, ORRUBA*, ...)



n arrays (*VANDLE. LENDA, NEXT, ODeSA*, ...)



Ge arrays (*GRET(IN)A, SeGA, Clovers,...)* 



### Constraining neutron-capture cross sections



- Neutron capture can occur via resonances (eg compound nucleus formation), and direct capture to bound states
- Depending on neutron level density/strong resonances, one or other may be dominant
- In general, need methods to constrain *both* mechanisms

#### Neutron transfer reactions

- Selective (states with target+n wavefunctions)
- Give properties of bound states and isolated resonances (E, J<sup>π</sup>, C<sup>2</sup>S)
  - for DSD
  - Constrain structure models
- Can be used as surrogate for CN capture (SRM)

### Constraining neutron-capture cross sections



- Neutron capture can occur via resonances (eg compound nucleus formation), and direct capture to bound states
- Depending on neutron level density, one or other may be
- Ideally need methods to constrain *both* mechanisms

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## Constraining n-capture cross sections - DSD



### Constraining CN n-capture cross sections - SRM



- Model reaction in HF formalism
- Essential theory components:
  - Formation of CN ( $\sigma^{CN}$ ) simple
  - Decay of CN (G<sup>CN</sup>) complicated (Escher)
  - Need to place experimental constraints on G<sup>CN</sup>

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A surrogate reaction forms the "same"\* compound nucleus as the desired reaction

- Model experimentally-determined  $P_{p\gamma}(E_{ex})$  in HF formalism to constrain  $G^{CN}$
- Essential theory components:
  - Decay of CN (*G<sup>CN</sup>*) complicated (Escher)
  - \*Entry spin distribution (FCN) complicated (Potel)

Figures and equations adapted from J.E. Escher et al. Rev. Mod. Phys. 84, 353 (2012).

## Constraining CN n-capture cross sections - SRM







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#### Ratkiewicz et al., Phys. Rev. Lett. 122, 052502 (2019)

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<sup>96</sup>Mo



#### Ratkiewicz et al., Phys. Rev. Lett. 122, 052502 (2019)

G. Potel

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## SRM outlook...

- SRM Progress
  - Validation of (d,p $\gamma$ ) as surrogate for (n,  $\gamma$ )
    - understanding  $\mathsf{J}^{\pi}$  formation distributions key ingredient
  - Development of (p,d) as (n,  $\gamma$ ) surrogate



- Ongoing development of (p,p') as  $(n, \gamma)$  surrogate
- What are the limits of the statistical approach...
- More from Jutta next week...

### Expt. challenges moving to RIB experiments

- Resolution (target thickness, kinematic compression) (~100 keV -> 500-1000 keV)
- Contaminants (eg carbon in target however, only need to address this 'once' – not a surprise every time)
- Luminosity (beam intensity) statistics-limited measurements
- Limitation on nuclides that can be practically studied
  - Nuclides without a reasonably-collecting transitions are very challenging - complex γ decay schemes disperses strength over numerous γ transitions
  - Isomers (moving with beam) lead to unobserved gamma emissions

An alternative technique – uniquely suited to FRIB...

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## A new technique with RIBs

...detect the recoiling nucleus to determine if decayed by n or  $\gamma$ Use charged-particle [e.g (d,p)] to determine the formation E<sub>x</sub>(as before)

*To bound states* <sup>84</sup>Se(d,p)<sup>85</sup>Se\*(γ)<sup>85</sup>Se

#### To unbound states

- ${}^{84}Se(d,p){}^{85}Se^{*}(\gamma){}^{85}Se$
- <sup>84</sup>Se(d,p)<sup>85</sup>Se\*(n)<sup>84</sup>Se







## A new technique with RIBs

...detect the recoiling nucleus to determine if decayed by n or  $\gamma$ Use charged-particle [e.g (d,p)] to determine the formation  $E_x$ (as before)

To bound states  $^{84}Se(d,p)^{85}Se^{*}(\gamma)^{85}Se$ 

#### To unbound states

- $^{84}Se(d,p)^{85}Se^{*}(\gamma)^{85}Se$
- <sup>84</sup>Se(d,p)<sup>85</sup>Se\*(n)<sup>84</sup>Se





et al.,

A1900



<sup>84</sup>Se

An alternative with RIBs...



H. Sims, J.A. Cizewski, S.D.Pain, A. Ratkiewicz,

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10-1

et al.,



### Advantages

- concentrates all statistics in a single observable
- high (25-50%), simple and experimentallydeterminable detection efficiency (cf  $\gamma$  cascades)
- Enables measurements on (almost) any nucleus
  on same footing

### Challenges

 need careful characterization of BG reactions on C in target

Work ongoing with Jutta to extract  $(n,\gamma)$  cross section

Excitation energy [MeV]

# Unique opportunity at FRIB

### Combination

- FRIB n-rich beams
- S800
- GODDESS [ORRUBA+GRET(IN)A]

### Two approved experiments

- <sup>80</sup>Ge (Sims, Grinder, Cizewski, Pain, et al)
   weak r process
- <sup>75</sup>Ga (Pain, Balakrishnan, et al)
  - i-process

Cannot measure all; target specific interesting cases

Ideally like to have (empirical) predictive model of  $(n,\gamma)$  cross sections

Model constrained by experiment in sensitive cases?
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### Detect protons, gammas and recoils

- Discrete particle-γ spectroscopy
- SRM with recoils
- SRM with  $\gamma$



## Transfer on sd nuclei

- Testing ground of LBSM calculations
  - Wealth of experimental data (near stability)
  - Well-constrained interactions for sd states (eg USDb)
  - Mid shell
    - highly mixed states
    - non-zero J<sup>π</sup> ground states
    - non-spherical systems
  - How well are *fp* excitations described?
- Reaction models
  - Lower-end of well-constrained nucleon-nucleus global potentials (near stability)

Na

16Ne

Jpi=0+ T1/2s-5.395E

**15F** Jpi=(1/2+) F1/2s=6.5821

140

13N

**12C** Jpi=0+ T1/2=>1.000E

- ADWA vs DWBA?
- Finite range effects?
- ...

- Beautiful experiments
- Astrophysical motivation N = Z

							Sc	36Sc	37Sc	38Sc	<b>39Sc</b> <sub>Jpi=(7/2-)</sub>	<b>40Sc</b> Jpi=4- T1/2s=1.823E	<b>41Sc</b> Jpi=7/2- T1/2s-5.963E	<b>42S</b> Jpi=0+ T1/3=-6.813E	<b>43Sc</b> Jpi=7/2- T1/2s-1.401E	<b>44Sc</b> Jpi=2+ T1/2s=1.414
5						Ca	34Ca <sub>Jpi-0+</sub>	<b>35Ca</b> T1/2s-5.000E	<b>36Ca</b> Jpi-0+ T1/2s=1.020E	<b>37Ca</b> Jpi-3/2+ T1/2s=1.811E	<b>38Ca</b> Jpi-0+ T1/2s=4.400E	<b>39Ca</b> Jpi-3/2+ T1/2s=8.596E	<b>40C*4</b> Jpi-0) T1/2×1.000E	<b>41Ca</b> Jpi-7/2- T1/2s=3.250E	<b>42Ca</b> Jpi-0+ T1/2s>1.000E	<b>43Ca</b> Jpi-7/2- T1/2s>1.000
					к	32K	33K	34K	<b>35K</b> Jpi=3/2+ T1/2s=1.900E	<b>36K</b> Jpi=2+ T1/2s=3.420E	<b>37K</b> Jpi=3/2+ T1/2s=1.226E	<b>38K</b> Jni=3+ T1/2.=4.582E	<b>39K</b> Jpi=3/2+ T1/2s>1.000E	<b>40K</b> Jpi=4- T1/2s=4.030E	<b>41K</b> Jpi=3/2+ T1/2s>1.000E	<b>42K</b> Jpi=2- T1/2s=4.450
es				Ar	30Ar	<b>31Ar</b> T1/2s=1.510E	<b>32Ar</b> Jpi=0+ T1/2s=9.800E	<b>33Ar</b> Jpi=1/2+ T1/2s=1.730E	<b>34Ar</b> Jpi=0+ T1/2s=8.445E	<b>35Ar</b> Jpi=3/2+ T1/2s=1.775E	<b>36A</b> J <sub>Di=0+</sub> T1/2~1.000E	<b>37Ar</b> Jpj=3/2+ T1/2s=3.027E	<b>38Ar</b> Jpi=0+ T1/2s>1.000E	<b>39Ar</b> J <sub>Di</sub> =7/2- T1/2s=8.489E	<b>40Ar</b> Jpj=0+ T1/2s>1.000E	<b>41Ar</b> Jpi=7/2- T1/2s=6.560
			CI	28CI	29CI	30CI	31CI T1/2s=1.500E	<b>32CI</b> Jpi=1+ T1/2s 2.980E	<b>33Cl</b> Jpi=3/2+ T1/2s 2.511E	<b>34C!</b> Jpi=0- T1/20 1.526E	<b>35CI</b> Jpi=3/2+ T1/2s>1.000E	<b>36CI</b> Jpi=2+ T1/2s=9.499E	<b>37CI</b> Jpi=3/2+ T1/2s>1.000E	<b>38CI</b> Jpi=2- T1/2s=2.234E	<b>39CI</b> Jpi=3/2+ T1/2s=3.336E	<b>40Cl</b> Jpi=2- T1/2s 8.100
		S	265	27S	<b>28S</b> Jpi=01 T1/2s=1.250H	<b>29S</b> Jpi=5/21 T1/2s=1.870E	<b>30S</b> Jpi=0+ T1/2s=1.178H	<b>31S</b> Jpi-1/2+ T1/2s=2.572H	328 Jpi-01 TT/2×1.000F	<b>33S</b> Jpi=3/2+ T1/2s>1.000E	<b>34S</b> Jpi=0+ T1/2s>1.000H	<b>35S</b> Jpi=3/21 T1/2s=7.561E	<b>36S</b> Jpi=01 T1/2s>1.000E	<b>37S</b> Jpi=7/2- T1/2s=3.030H	<b>38S</b> Jpi=01 T1/2s=1.022H	<b>39S</b> Jpi=(3/2.5/2 T1/2s=1.150
	Ρ	24P	25P	<b>26P</b> Jpi=(3+) T1/2s=2.000E	<b>27P</b> Jpi=1/2+ T1/2s=2.600E	<b>28P</b> Jpi=3+ T1/2s=2.703E	<b>29P</b> Jpi=1/2+ T1/2s=4.140E	<b>30P</b> Jni=1+ T1/2 =1.499E	<b>31P</b> Jpi=1/2+ T1/2s>1.000E	<b>32P</b> Jpi=1+ T1/2s=1.232E	<b>33P</b> Jpi=1/2+ T1/2s=2.189E	<b>34P</b> Jpi=1+ T1/2s=1.243E	<b>35P</b> Jpi=1/2+ T1/2s=4.730E	<b>36P</b> T1/2s=5.600H	<b>37P</b> T1/2s=2.310E	<b>38P</b> T1/2s=6.400
Si	<b>22Si</b> Jpi=0+ T1/2s=6.000E	23Si	<b>24Si</b> Jpi=0+ T1/2s=1.020E	<b>25Si</b> Jpi=5/2+ T1/2s=2.200E	<b>26Si</b> Jpi=0+ T1/2s=2.234E	<b>27Si</b> <sub>Jpi=5/2+</sub> T1/2s=4.160E	285 Jpi=0+ T1/2.>1.000E	<b>29Si</b> J <sub>Di=1/2+</sub> T1/2s>1.000E	<b>30Si</b> J <sub>DI=0+</sub> T1/2∽1.000E	<b>31Si</b> <sub>Jpi=3/2+</sub> T1/2s-9.438E	<b>32Si</b> Jpi=0+ T1/2s=5.428E	<b>33Si</b> T1/2s=6.180E	<b>34Si</b> Jpi=0+ T1/2s=2.770E	<b>35Si</b> T1/2s=7.800E	<b>36Si</b> <sub>Jpi=0+</sub> T1/2s=4.500E	37Si
AI	21AI	<b>22AI</b> T1/2s-7.000E	<b>23AI</b> T1/2s-4.700E	<b>24AI</b> Jpi-4+ T1/2s=2.053E	<b>25AI</b> Jpi-5/2+ T1/2s=7.183E	<b>26A!</b> Jpi-5+ T1/2-2.335E	<b>27AI</b> Jpi-5/2+ T1/2s>1.000E	<b>28AI</b> <sup>Jpi-3+</sup> T1/2s=1.345E	<b>29AI</b> Jpi-5/2+ T1/2s=3.936E	<b>30AI</b> Jpi-3+ T1/2s=3.600E	<b>31AI</b> Jpi-(3/2,5/2)+ T1/2s=6.440E	<b>32AI</b> Jpi-1+ T1/2s=3.300E	33AI	<b>34AI</b> T1/2s-6.000E	<b>35AI</b> T1/2s-1.500E	36AI
Mg	<b>20Mg</b> Jpi=0+ T1/2s=9.500E	<b>21Mg</b> Jpi=(3/2,5/2)+ T1/2s=1.220E	<b>22Mg</b> Jpi=0+ T1/2s=3.857E	<b>23Mg</b> Jpi=3/2+ T1/2s=1.132E	24M Jpi=0+ T1/2=1.000E	<b>25Mg</b> Jpi=5/2+ T1/2s>1.000E	<b>26Mg</b> Jpi=0+ T1/2s>1.000E	<b>27Mg</b> Jpi=1/2+ T1/2s=5.675E	<b>28Mg</b> Jpi=0+ T1/2s=7.528E	<b>29Mg</b> Jpi=3/2+ T1/2s=1.300E	<b>30Mg</b> Jpi=0+ T1/2s=3.350E	31Mg T1/2s=2.300E	<b>32Mg</b> Jpi=0+ T1/2s=1.200E	33Mg T1/2s=9.000E	<b>34Mg</b> Jpi=0+ T1/2s=2.000E	35Mg
18Na	19Na	<b>20Na</b> <sub>Jpi=2+</sub> T1/2s=4.479E	<b>21Na</b> Jpi=3/2+ T1/2s=2.249E	<b>22N</b> Jpi=3+ T1/2 =8.211E	<b>23Na</b> <sub>Jpj=3/2+</sub> T1/2s>1.000E	<b>24Na</b> Jpi=4+ T1/2s=5.385E	<b>25Na</b> Jpi=5/2+ T1/2s=5.910E	<b>26Na</b> <sub>Jpj=3+</sub> Tl/2s=1.072E	<b>27Na</b> Jpi=5/2+ T1/2s=3.010E	<b>28Na</b> Jpi=1+ T1/2s=3.050E	<b>29Na</b> Joi=3/2 T1/2s=4.490E	<b>30Na</b> J <sub>Dj=2+</sub> T1/2s=4.800E	<b>31Na</b> Jpi=3/2+ T1/2s=1.700E	<b>32Na</b> Jpi=(3-,4-) T1/2s=1.320E	<b>33Na</b> T1/2s=8.200E	34Na T1/2s=5.500
<b>17Ne</b> Jpi=1/2- T1/2s-1.092E	<b>18Ne</b> Jpi=0+ T1/2s-1.672E	<b>19Ne</b> Jpi=1/2+ T1/2s=1.722E	20N Jpi=0+ T1/2 >1.000E	<b>21Ne</b> Jpi=3/2+ T1/2s>1.000E	<b>22Ne</b> Jpi=0+ T1/2s>1.000E	<b>23Ne</b> Jpi=5/2+ T1/2s=3.724E	<b>24Ne</b> Jpi=0+ T1/2s-2.028E	25Ne Jpi=(1/2,3/2)+ T1/2s-6.020E	<b>26Ne</b> Jpi=0+ T1/2s-1.970E	27Ne T1/2s=3.200E	28Ne Jpi=0+ T1/2s=1.700E	<b>29Ne</b> T1/2s=2.000E	<b>30Ne</b> <sub>Jpi=0+</sub>	31Ne	<b>32Ne</b> Jpi=0+	23
<b>16F</b> Jpi-0- T1/2s=1.645E	<b>17F</b> Jpi=5/2+ T1/2s=6.449B	<b>18F</b> Jpi=11 T1/2 =6.586F	<b>19F</b> Jpi-1/21 T1/2s>1.000E	<b>20F</b> Jpi=21 T1/2s=1.100E	<b>21F</b> Jpi-5/21 T1/2s=4.158H	<b>22F</b> Jpi=41,(31) T1/2s=4.230E	<b>23F</b> Jpi=(3/2,5/2)+ T1/2s=2.230E	<b>24F</b> Jpi=(1,2,3)1 T1/2s=3.400E	25F T1/2s=5.900E	26F	27F	28F	29F	21	22	
<b>150</b> Jpi=1/2- T1/2s=1.222E	160 Jpi=0+ T1/2 > 1.000E	<b>170</b> Jpi=5/2+ T1/2s>1,000E	<b>180</b> Jpi=0+ T1/2s>1.000E	<b>190</b> Jpi=5/2+ T1/2s=2.691E	<b>200</b> Jpi=0+ T1/2s=1.351E	<b>210</b> Jpi=(1/2,3/2,5, T1/2s=3,420E	<b>220</b> Jpi=0+ T1/2s=2.250E	<b>230</b> T1/2s=8.200F/	<b>240</b> Jpi=0+ T1/2s=6.100E	250	260	19	20	]		
<b>14N</b> Joi=1+ T1/2->1.000E	<b>15N</b> Jpi=1/2- T1/2s>1.000E	<b>16N</b> Jpi=2- T1/2s=7.130E	<b>17N</b> Jpi=1/2- T1/2s=4.173E	<b>18N</b> Jpi=1- T1/2s=6.240E	<b>19N</b> Jpi=(1/2-) T1/2s=3.040E	<b>20N</b> T1/2s=1.000E	<b>21N</b> T1/2s=8.500E	<b>22N</b> T1/2s=2.400E	23N	24N	18					
<b>13C</b> Jpi-1/2- T1/2s>1.000E	<b>14C</b> Jpi-0+ T1/2s=1.808E	<b>15C</b> Jpi=1/2+ T1/2s=2.449E	<b>16C</b> Jpi-0+ T1/2s=7.470E	<b>17C</b> T1/2s-1.930E	<b>18C</b> Jpi-0+ T1/2s=9.500E	<b>19C</b> T1/2s-4.600E	<b>20C</b> Jpi-0+ T1/2s=1.400E	21C	<b>22C</b> <sub>Jpi-0+</sub>	17	]					

## Odd-odd N=Z sd-shell nuclides

N=Z

IS

Na

16Ne

15F

140

13N

16F

14 N

- Networks of  $(p,\gamma)$  and  $(p,\alpha)$  reactions (beta decays omitted) in novae
- $(p,\gamma)$  on odd-odd N = Z nuclides particularly important
  - bottleneck reactions

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- impact on astronomical observables
- Many have long-lived spin isomers that can play roles in reaction networks (astromers)

S	pecies	t(Gs)	†(ıs)	J <sup>π</sup> (GS)	J <sup>π</sup> (IS)	
_	<sup>22</sup> Na	2.6 y	240 ns	3+	1+	
_	<sup>24</sup> Al	2 s	130 ms	4+	1+	
_	<sup>26</sup> Al	0.7 My	6.3 s	5+	0+	
_	<sup>30</sup> P	2.5 m	96 fs	1+	0+	
_	<sup>34</sup> Cl	1.5 s	32 m	0+	3+	
_	<sup>38</sup> K	6.7 m	0.9 s	3+	0+	
_	<sup>42</sup> Sc	0.7 s	1 m	0+	7+	



# Odd-odd N=Z nuclides - isomers

N=Z

Na

16Ne

1/2s=5.395E

15F

140

13N

- Networks of (p,γ) and (p,α) reactions (beta decays omitted) in novae
- (p, $\gamma$ ) on odd-odd N = Z nuclides
  - bottleneck reactions
  - impact on astronomical observables
- Reaction networks
  - independent
  - or thermal coupling at high T
- Want reaction rates on both GS <sup>IS</sup> and IS
  - v. different SP structure, limited expt
  - sdpf states
- General rule
  - Insufficient beam intensities for direct (p,γ) measurements currently (some at FRIB)
- indirect techniques within reach
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## Transfer reactions on isomeric beams at FRIB

- Unique design of FRIB gives opportunities for producing beams of long-lived (>ms) nuclear isomers at ideal beam energies/optics for transfer reactions - ReA
- Produce GS and IS with fragmentation, reaccelerate
- Control the ground:isomer composition via
  - selection of production yields, via fragment separator (spin, though not specifically that of final state)
  - Adjustment of hold-up times inherent to ReA (lifetimes)

- Transfer, charge-exchange, Coulex,... for structure/indirect astrophysics (ORRUBA, SOLARIS, GODDESS, GRETA, LENDA, SeGA, Clarion2, ....)
- Direct measurements of astrophysical reactions [eg (p,g) with SECAR, or (a,p) (a,n) with JENSA, MUSIC, HabaNERO, ...]



### Previous and upcoming isomer expts

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*Nova nucleosynthesis* <sup>34g,m</sup>Cl(d,p) FRIB PAC1 approved



# Indirect constraints on $(p,\gamma)$ reactions

- Dominated by isolated resonances (some too low for direct measurements)
- Locate states E<sub>r</sub>
- To constrain resonance strength, determine:
  - Spins
    - $\ell_{\rho}$  (constrains barrier,  $\Gamma_{sp}$ )
  - Determine reduced width (10<sup>2</sup>)  $\Gamma_p = C^2 S. \Gamma_{sp}$





1e-06

1e-07

200

400

 $E_{u}$  (keV)

600

 $\Gamma_{v}(E1)$ 

 $--- \Gamma_{v}(M1)$ 

1000

# What can we learn from transfer?

- Proton transfer ideal (d,n) or (<sup>3</sup>He,d)
  - Selectivity
  - Energies 10s of keV
  - *l*<sub>p</sub>
  - *C*<sup>2</sup>*S*
- Experimental challenges
  - Neutron detection
  - Inclusive measurements
    - Gamma tagging
    - Non-spin-zero ground/isomeric states

 $\Gamma_p = C^2 S. \Gamma_{sp}$ 

 $\omega \gamma \approx \omega \Gamma_p$ 

- <sup>3</sup>He targets
- Infer C<sup>2</sup>S for *single-proton states* via *mirror symmetry*   $C^{2}S_{p} \approx C^{2}S_{n}$ 
  - Guide by SMEC
  - 10-20% effect







<sup>26</sup>Al(d,p)<sup>27</sup>Al experiment



- 4.5 MeV/u <sup>26</sup>Al (Oak Ridge Tandem)
- 5x10<sup>6</sup> pps
- 150 μg/cm<sup>2</sup> CD<sub>2</sub>
- MCP normalization (200 kHz)
- •Large Q value = low kinematic compression







<sup>26</sup>Al *sd*-shell states

#### What to expect?



*fp*-shell excitations at higher E<sub>x</sub>

 $J_t - j \leq J_f \leq J_t + j$ 





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#### What to expect?

- sd-shell excitations
- *fp*-shell excitations at higher E<sub>x</sub>

 $J_t - j \leq J_f \leq J_t + j$ 

<sup>26</sup>AI(d,p)<sup>27</sup>AI angular distributions

### total

#### s wave

p wave

#### d wave fwave

#### Important for quality analysis of transfer data

- Good coverage of first stripping peak
- ADWA
- Finite Range (d and T)
- KD nucleon-nucleus **OMPs**
- Standard geometry parameters









## Comparison with USDB - energies

Using NuShellX@MSU

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- Calculated the first 40 states of each positive parity (5/2+ -17/2+) in <sup>27</sup>Al (dim 80,000) and calculate overlaps with <sup>26</sup>Al 5+ ground state (dim 70,000) (few minutes with 6core/12-threads (280 states total)
- Match by comparing C<sup>2</sup>S and energy



#### **Excellent agreement in energies**

- Perhaps unsurprising, as USDb fitted to energy levels in the SD shell (608 states, 77 nuclei, incl <sup>26,27</sup>Al)
  - Only fitted up until J<sup>π</sup> becomes ambiguous - typically around 6-7 MeV
  - Agreement up to 11 MeV in expt



FIG. 2. (Color) Number of states used for the USDA and USDB Hamiltonians for each nucleus.

## Comparison with USDb



Macfarlane & French sum rules

$$N_h = \sum_{E} \frac{2J_f + 1}{2J_t + 1} C^2 S_{(d,p)}$$

$$N_n + N_h = 12$$
 (for sd shell)

- O Vacancies 7
- Occupancies 5
- O sd-shell space 12





# Summary

- Impossibly broad scope of FRIB transfer reaction program - focus on a couple of unique opportunities
- Developments in SRM: (d,p) for  $(n,\gamma)$
- Opportunities for (d,p) measurements at the \$800
  - strong beams in the mass ~80 region
  - discrete particle and particle-γ spectroscopy
  - New technique: SRM for CN neutron transfer using recoils, and  $\gamma$ s
  - Two approved experiments GODDESS+S800 upcoming
- Ability to produce high-quality Reacc. beams with controllable isomer content
  - Mirror studies for (p,  $\gamma$ ) reactions on isomers, ultimately (p,  $\gamma$ ) SECAR
  - Direct ( $\alpha$ ,p) measurements (JENSA) on isomers, ultimately SECAR
  - Two approved ReA experiments <sup>34</sup>Cl(d,p) and <sup>26</sup>Al/<sup>26</sup>Si(a,p)
  - expt/SM across the sd N=Z ( $^{22}Na$ ,  $^{30}P$ ,  $^{34g,m}Cl$ ,  $^{26g,m}Al$ ,  $^{38g,m}K$ )









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