

# Microscopic optical potentials for the FRIB era

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Texas A&M University, College Station

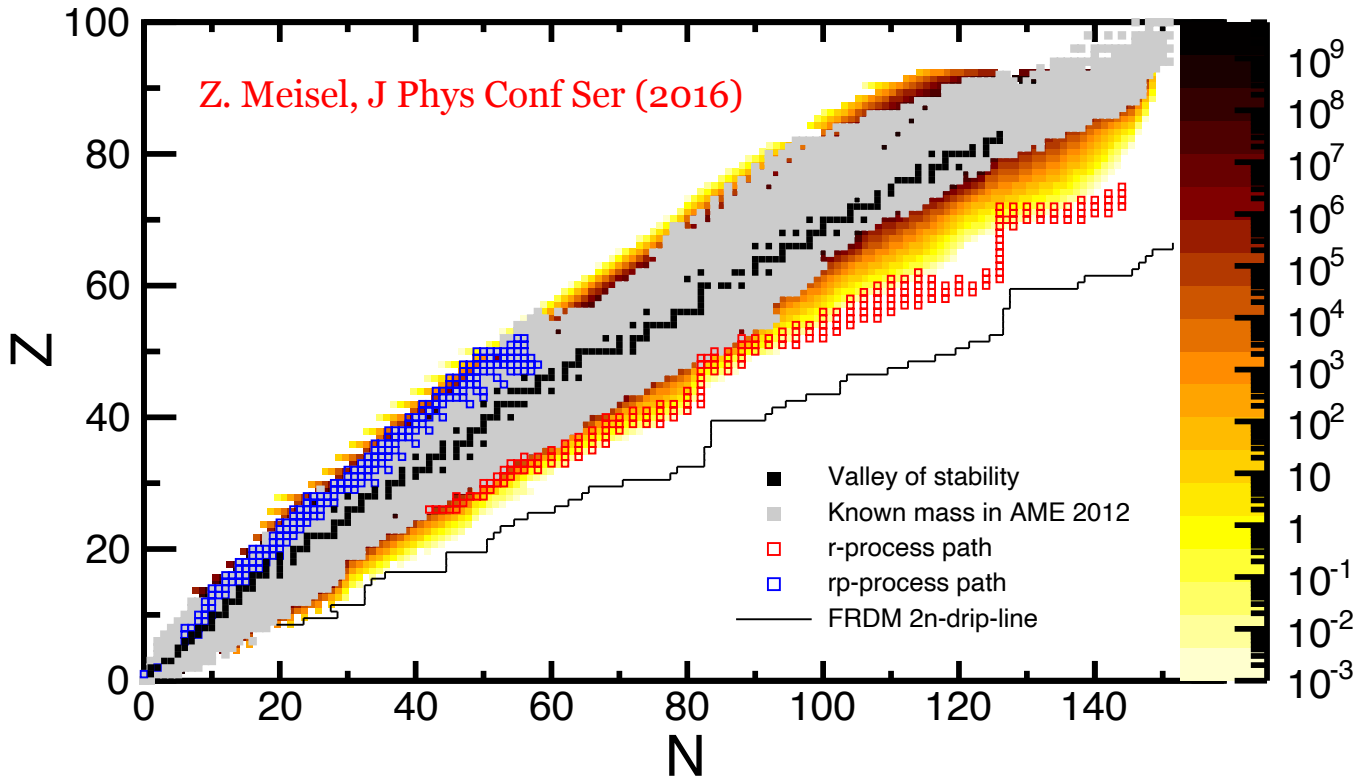
\*Collaborators: Taylor Whitehead, Yeunhwan Lim, Pengsheng Wen



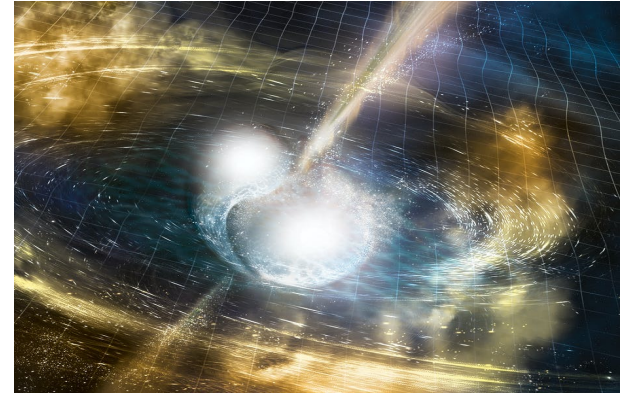
## Motivation: Nuclear reactions with unstable isotopes

- Interpretation of rare isotope beam experiments (neutron skins, nuclear properties near the dripline, capture rates in astrophysical environments)
  
- Challenges:
  - Ab initio nuclear reaction theory available for light systems at low energies
  - *Optical potentials* are a necessary ingredient in most reaction models, including transport analysis of heavy-ion collisions
  - *Global* optical potentials are needed for a consistent reaction theory over a wide range of unstable isotopes
  - Commonly used global optical potentials are phenomenological and fitted around nuclear stability
  
- Advantages of *microscopic approaches*
  - Start from the same many-body principles governing ab initio nuclear structure theory
  - Can reveal more efficient parameterizations for phenomenological optical potentials
  - Can provide a suitable *prior* as part of a more comprehensive Bayesian uncertainty analysis that incorporates scattering data

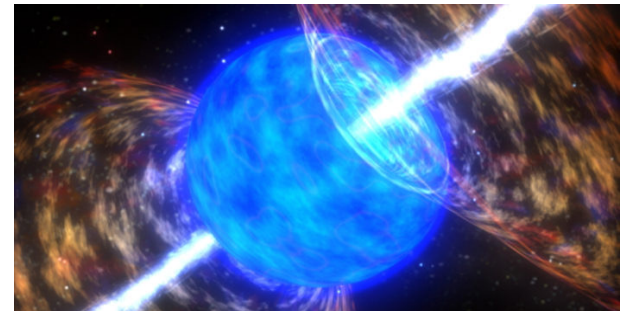
# Application: R-process nucleosynthesis



Neutron star mergers



Astrophysical site?

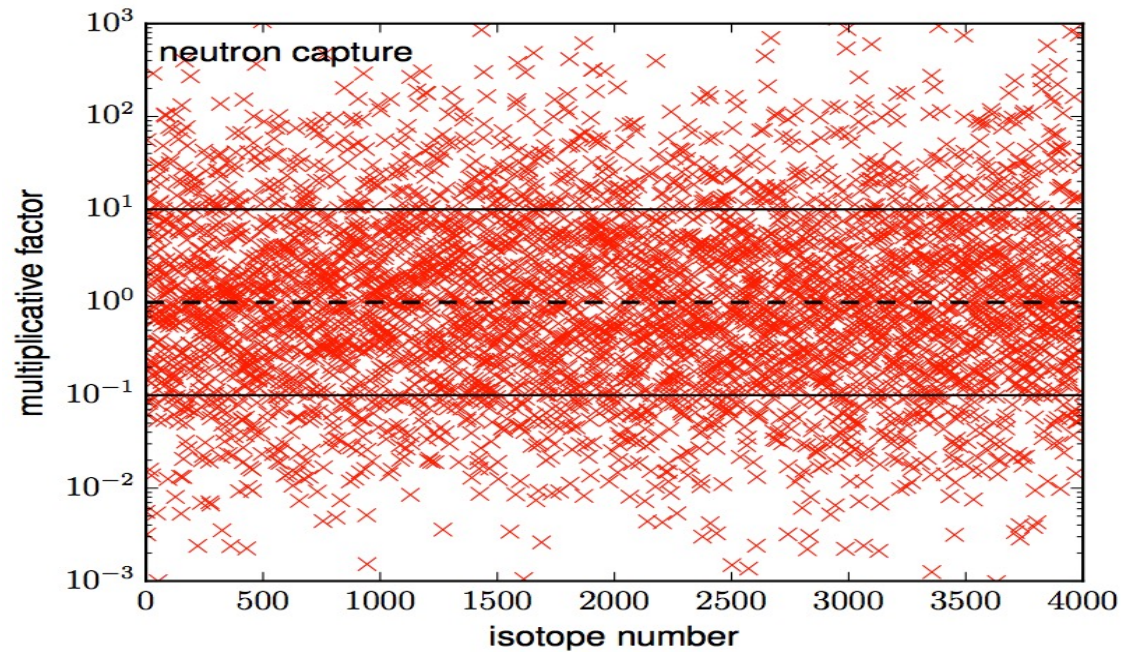


Collapsars

# Neutron capture sensitivity studies

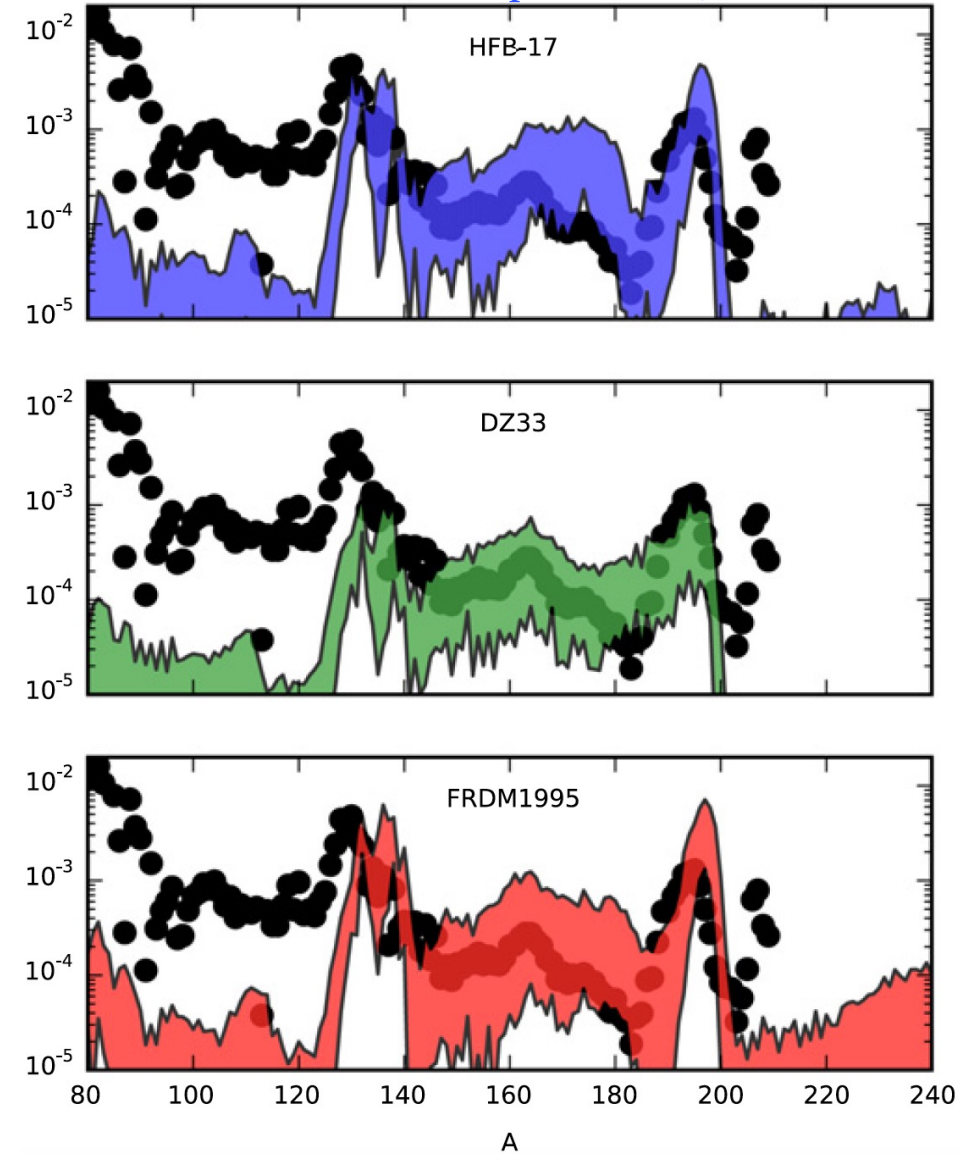
## Uncertainties coming from:

- ▶ Neutron-nucleus optical potentials
- ▶ Nuclear level densities for Hauser-Feshbach
- ▶  $\gamma$  strength functions



Mumpower et al., JPCF (2015)

Mumpower et al., PPNP (2016)

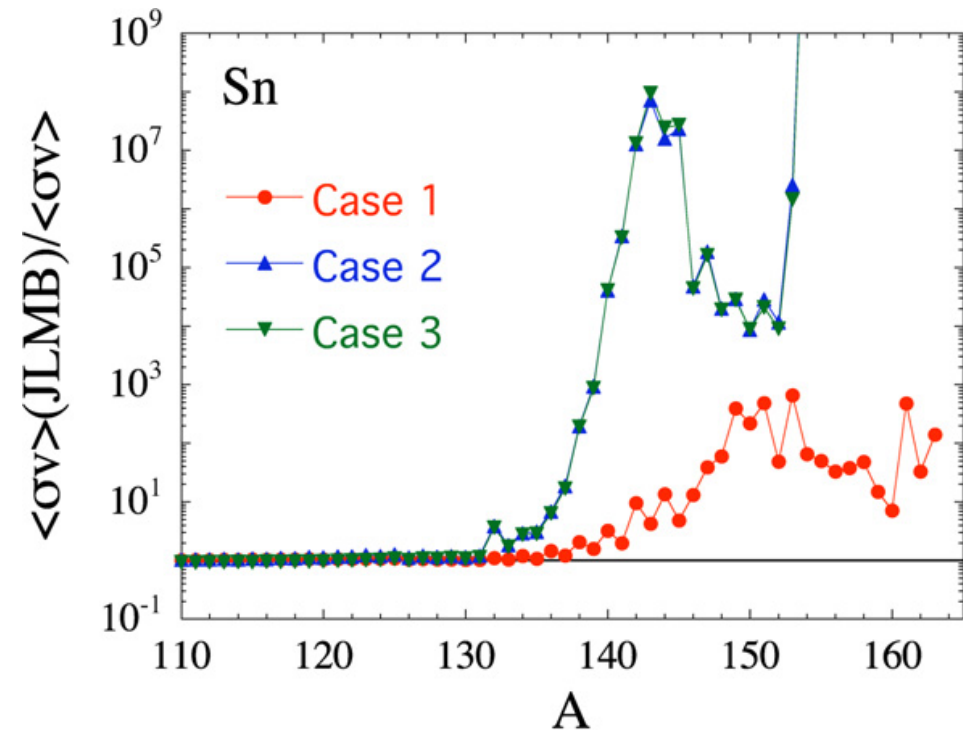
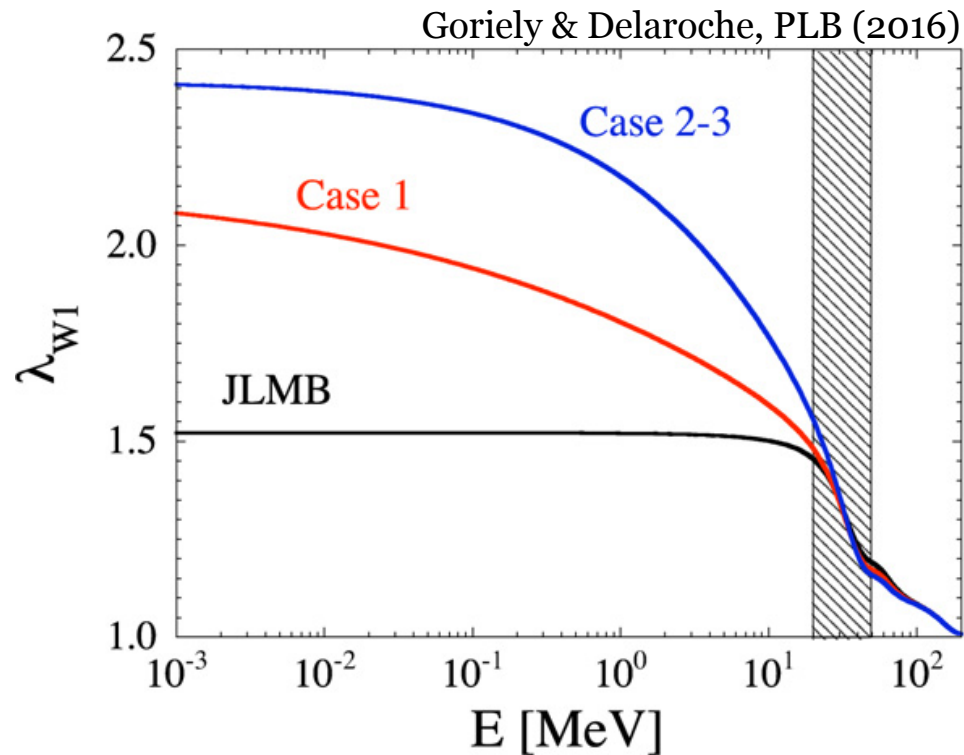


## Uncertainties coming from:

- ▶ Neutron-nucleus optical potentials

$$U(E) = \lambda_V(E) [V_0(E) + \lambda_{V_1}(E) \alpha V_1(E)] + i \lambda_W(E) [W_0(E) + \lambda_{W_1}(E) \alpha W_1(E)]$$

$\alpha = (\rho_n - \rho_p) / \rho$

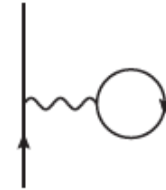


# Global microscopic optical potential: start from nucleon optical potential in infinite nuclear matter

- Identified with the on-shell nucleon self-energy  $\Sigma(\vec{r}_1, \vec{r}_2, \omega)$

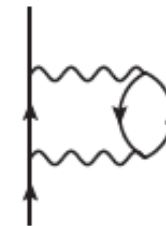
- First-order (Hartree-Fock) contribution (real & energy independent):

$$\Sigma_{2N}^{(1)}(q; k_f) = \sum_1 \langle \vec{q} \vec{h}_1 s s_1 t t_1 | \bar{V}_{2N} | \vec{q} \vec{h}_1 s s_1 t t_1 \rangle n_1$$



- Second-order perturbative contribution (complex & energy dependent):

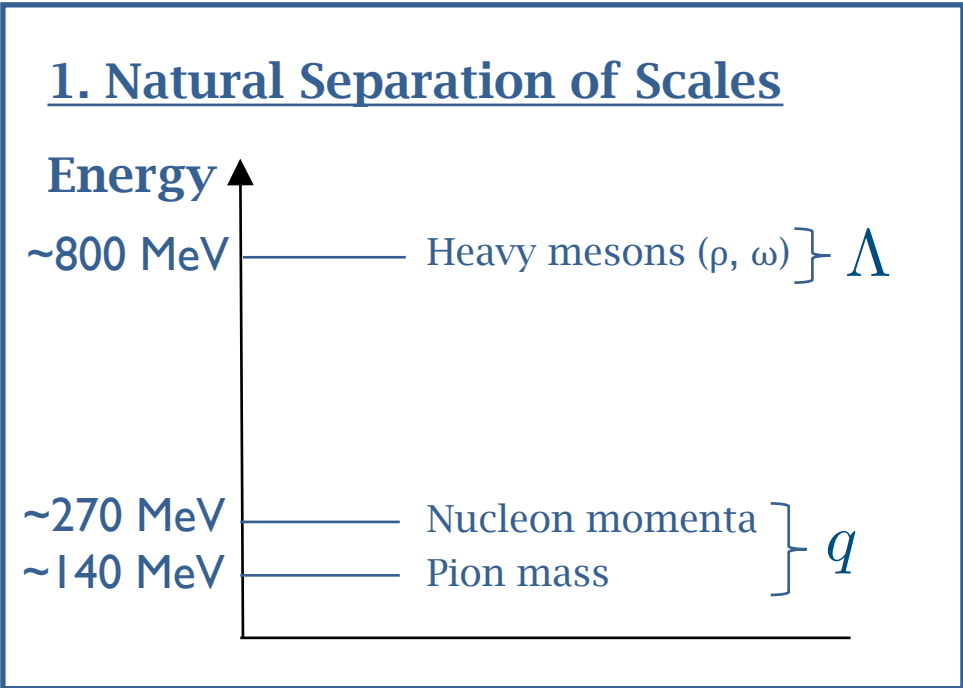
$$\Sigma_{2N}^{(2a)}(q, \omega; k_f) = \frac{1}{2} \sum_{123} \frac{|\langle \vec{p}_1 \vec{p}_3 s_1 s_3 t_1 t_3 | \bar{V} | \vec{q} \vec{h}_2 s s_2 t t_2 \rangle|^2}{\omega + \epsilon_2 - \epsilon_1 - \epsilon_3 + i\eta} \bar{n}_1 n_2 \bar{n}_3$$



- Self consistency:  $\epsilon(q) = \frac{q^2}{2M} + \text{Re} \Sigma(q, \epsilon(q))$

- Derived from high-precision chiral two- and three-body forces

# Chiral effective field theory (EFT) for nuclear forces



### 2. Goldstone bosons (pions) weakly-coupled at low momenta

$$\mathcal{L}_{\pi\pi}^{(2)} = \frac{1}{2} \partial_\mu \vec{\pi} \cdot \partial^\mu \vec{\pi} + \frac{1}{2f_\pi^2} (\partial_\mu \vec{\pi} \cdot \vec{\pi})^2$$

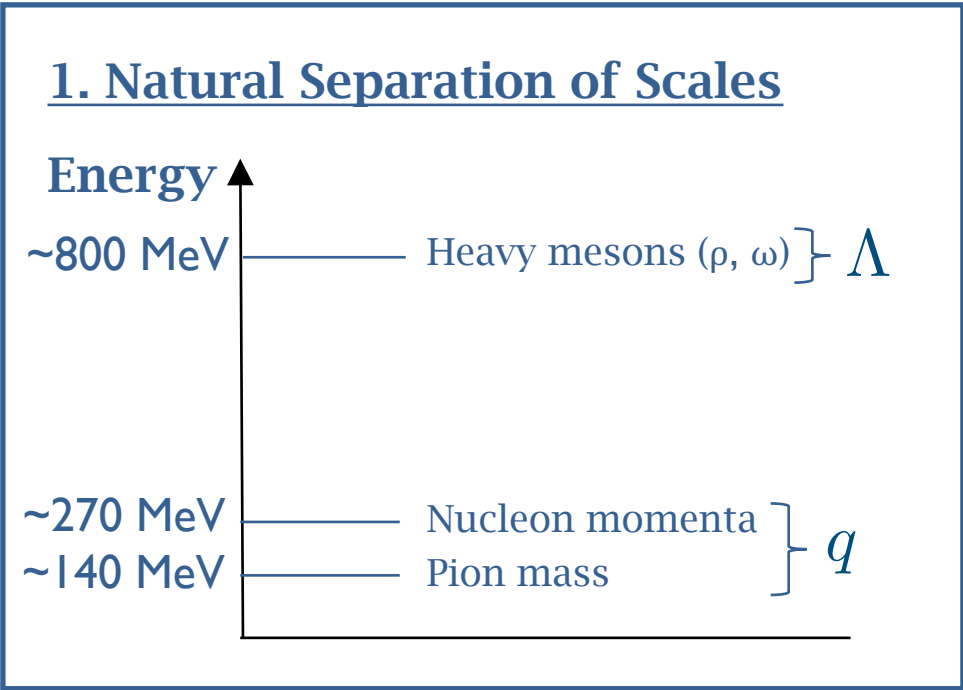
$$\mathcal{L}_{\pi N}^{(1)} = \bar{N} \left( i\gamma^\mu D_\mu - m - \frac{g_A}{2f_\pi} \gamma^\mu \gamma_5 \vec{\tau} \cdot \partial_\mu \vec{\pi} \right) N$$

## Nuclear forces from chiral EFT

	2N force	3N force	4N force
$(q/\Lambda)^0$		Systematic expansion	
$(q/\Lambda)^2$			
$(q/\Lambda)^3$			
$(q/\Lambda)^4$			

Enables uncertainty quantification

# Chiral effective field theory (EFT) for nuclear forces

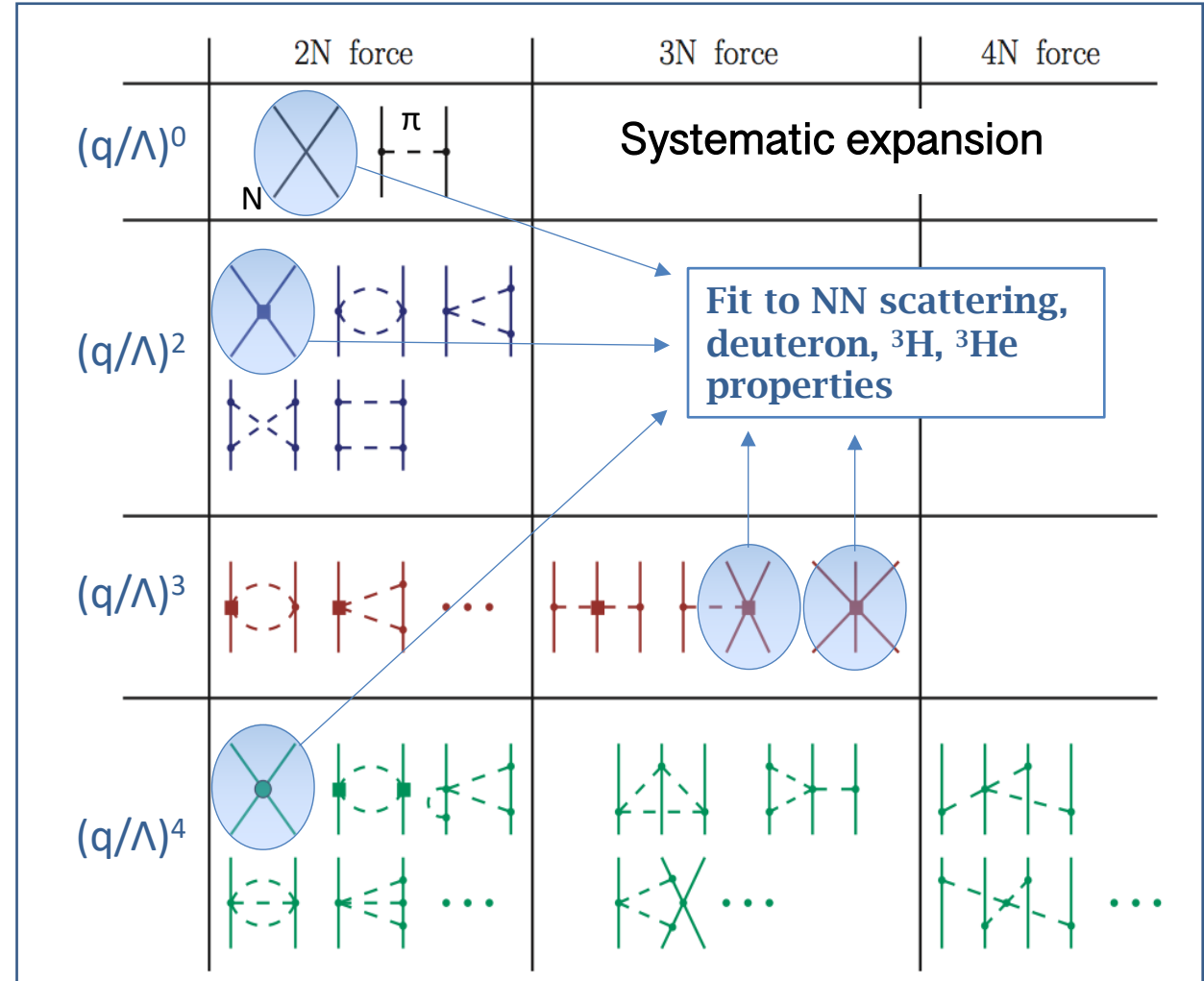


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## Nuclear forces from chiral EFT



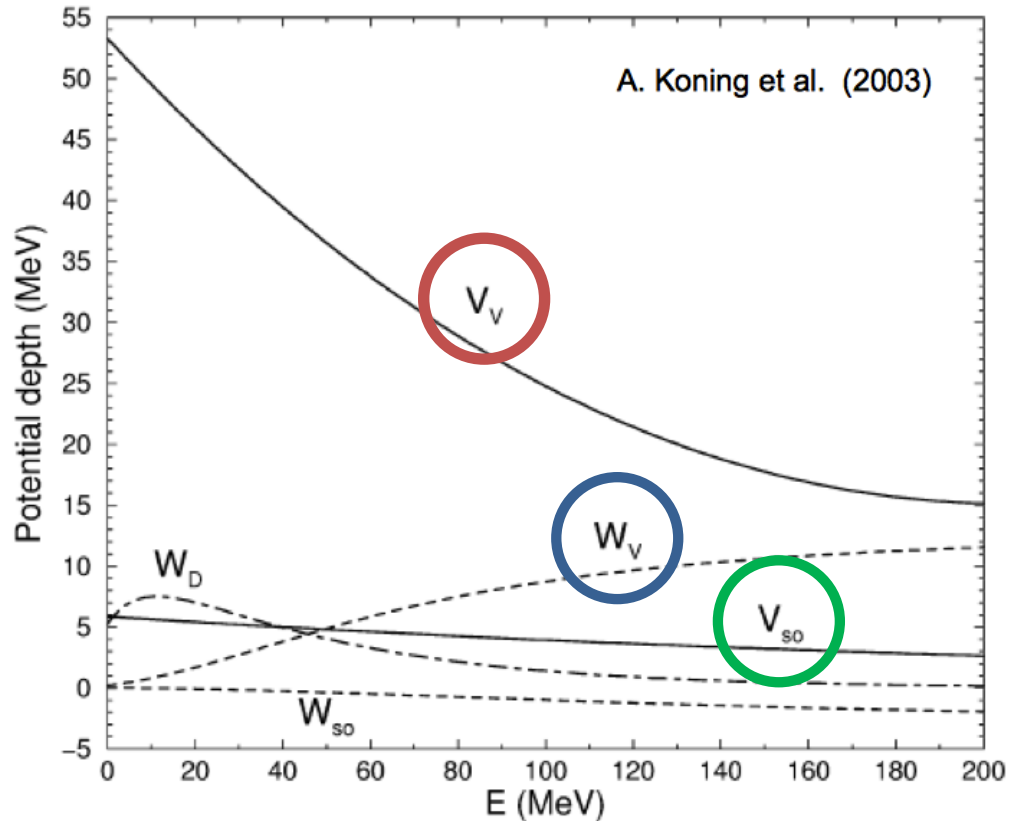
Enables uncertainty quantification



# Global optical potential parameterization



$$\mathcal{U}(r, E) = -\mathcal{V}_V(r, E) - i\mathcal{W}_V(r, E) - i\mathcal{W}_D(r, E) + \mathcal{V}_{SO}(r, E) \cdot \mathbf{l} \cdot \boldsymbol{\sigma} + i\mathcal{W}_{SO}(r, E) \cdot \mathbf{l} \cdot \boldsymbol{\sigma} + \mathcal{V}_C(r).$$



Woods-Saxon shape:  $f(r, R_i, a_i) = (1 + \exp[(r - R_i)/a_i])^{-1}$

$$\mathcal{V}_V(r, E) = V_V(E) f(r, R_V, a_V),$$

$$\mathcal{W}_V(r, E) = W_V(E) f(r, R_V, a_V),$$

$$\mathcal{W}_D(r, E) = -4a_D W_D(E) \frac{d}{dr} f(r, R_D, a_D),$$

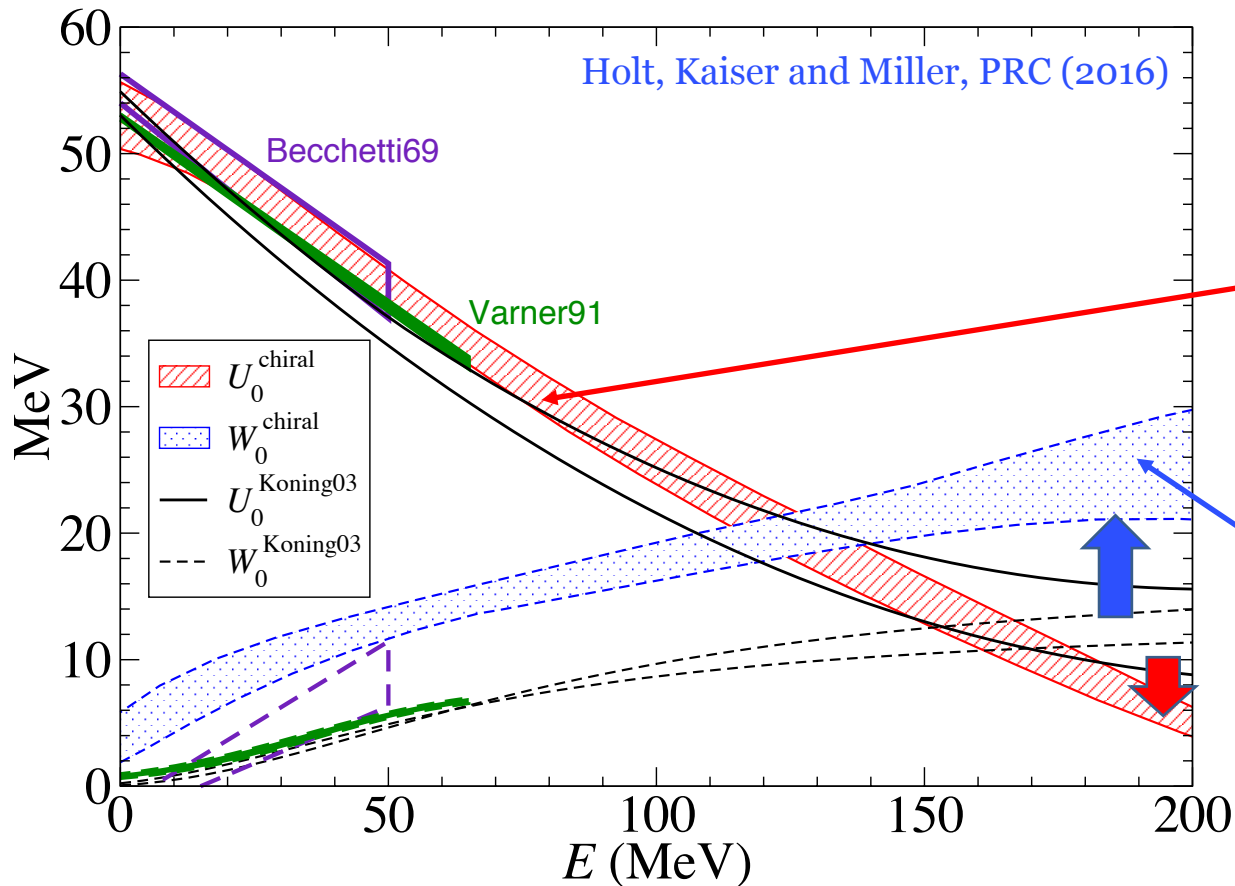
$$\mathcal{V}_{SO}(r, E) = V_{SO}(E) \left( \frac{\hbar}{m_\pi c} \right)^2 \frac{1}{r} \frac{d}{dr} f(r, R_{SO}, a_{SO}),$$

$$\mathcal{W}_{SO}(r, E) = W_{SO}(E) \left( \frac{\hbar}{m_\pi c} \right)^2 \frac{1}{r} \frac{d}{dr} f(r, R_{SO}, a_{SO}).$$

# Global optical potential parameterization: volume terms



$$\begin{aligned}
 \mathcal{U}(r, E) = & -\mathcal{V}_V(r, E) - i\mathcal{W}_V(r, E) - i\mathcal{W}_D(r, E) \\
 & + \mathcal{V}_{SO}(r, E) \cdot \mathbf{l} \cdot \boldsymbol{\sigma} + i\mathcal{W}_{SO}(r, E) \cdot \mathbf{l} \cdot \boldsymbol{\sigma} + \mathcal{V}_C(r).
 \end{aligned}$$



$$\mathcal{V}_V(r, E) = V_V(E) f(r, R_V, a_V),$$

Real volume strength function:

Very good agreement with phenomenology

$$\mathcal{W}_V(r, E) = W_V(E) f(r, R_V, a_V)$$

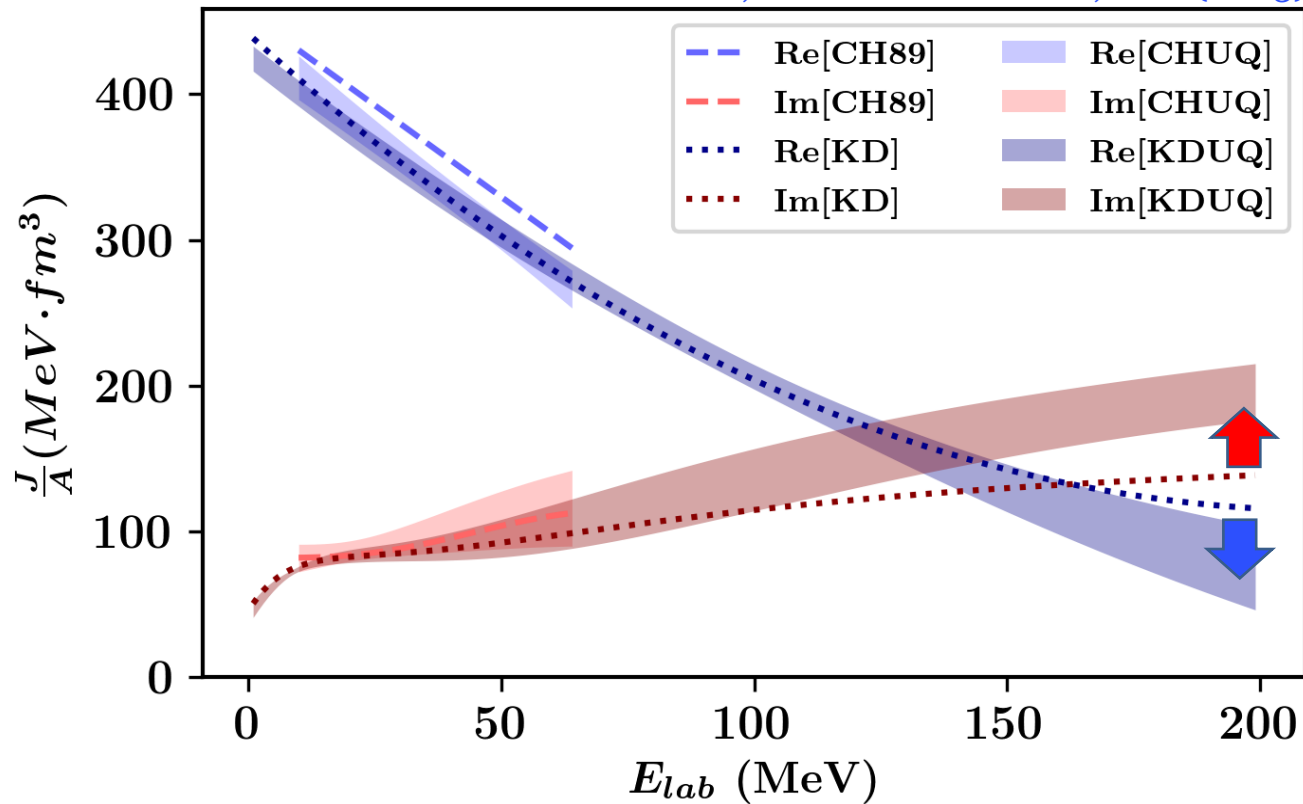
Imaginary volume strength function:

Too strongly absorptive

# Global optical potential parameterization: volume terms

$$\mathcal{U}(r, E) = -\mathcal{V}_V(r, E) - i\mathcal{W}_V(r, E) - i\mathcal{W}_D(r, E) + \mathcal{V}_{SO}(r, E) \cdot \mathbf{l} \cdot \boldsymbol{\sigma} + i\mathcal{W}_{SO}(r, E) \cdot \mathbf{l} \cdot \boldsymbol{\sigma} + \mathcal{V}_C(r).$$

Pruitt, Escher and Rahman, PRC (2023)



$$\mathcal{W}_V(r, E) = \boxed{W_V(E)} f(r, R_V, a_V)$$

Imaginary volume strength increases

$$\mathcal{V}_V(r, E) = \boxed{V_V(E)} f(r, R_V, a_V),$$

Real volume strength decreases

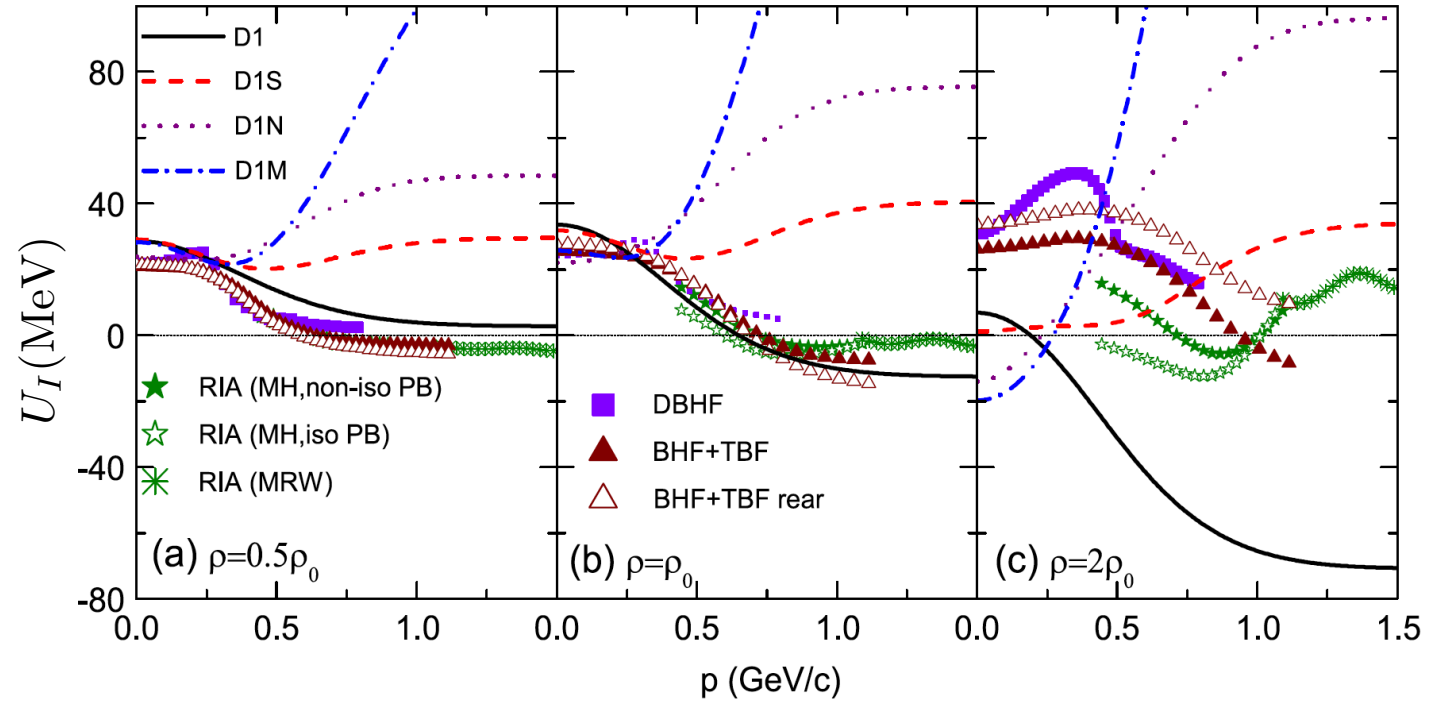
# Optical potentials in neutron-rich infinite matter

- Isovector optical potential obeys “Lane form”:

$$U_V = U_0 - U_I \delta_{np} \tau_z$$

- Much larger uncertainties for the isovector optical potential
- Isospin inversion ( $U_p > U_n$ ) at high energies present in only some models

B.-A. Li et al., PPNP (2018)



- Isovector Lane form above not obeyed at high energy in phenomenological Koning-Delaroche optical potential

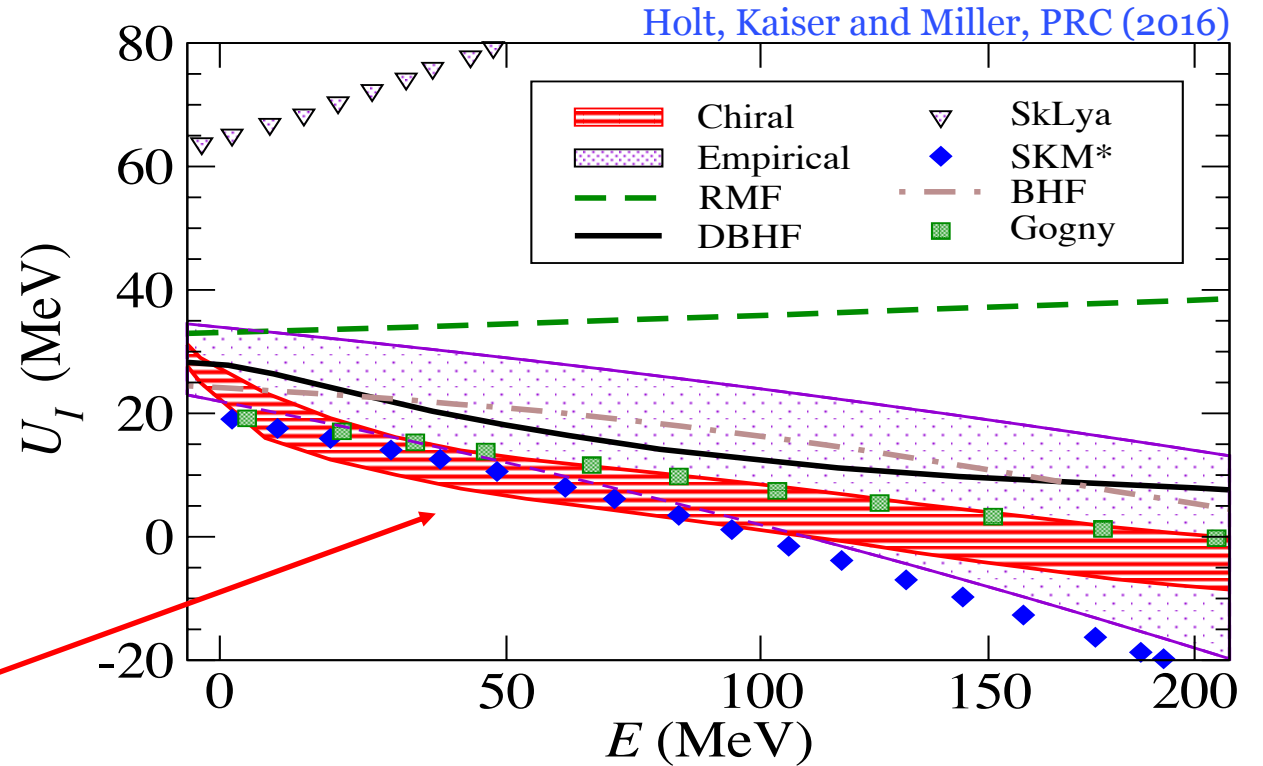
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- Chiral effective field theory:

- Matches energy dependence of phenomenological potentials well
- Predicts isospin inversion around  $E \approx 150 - 200$  MeV

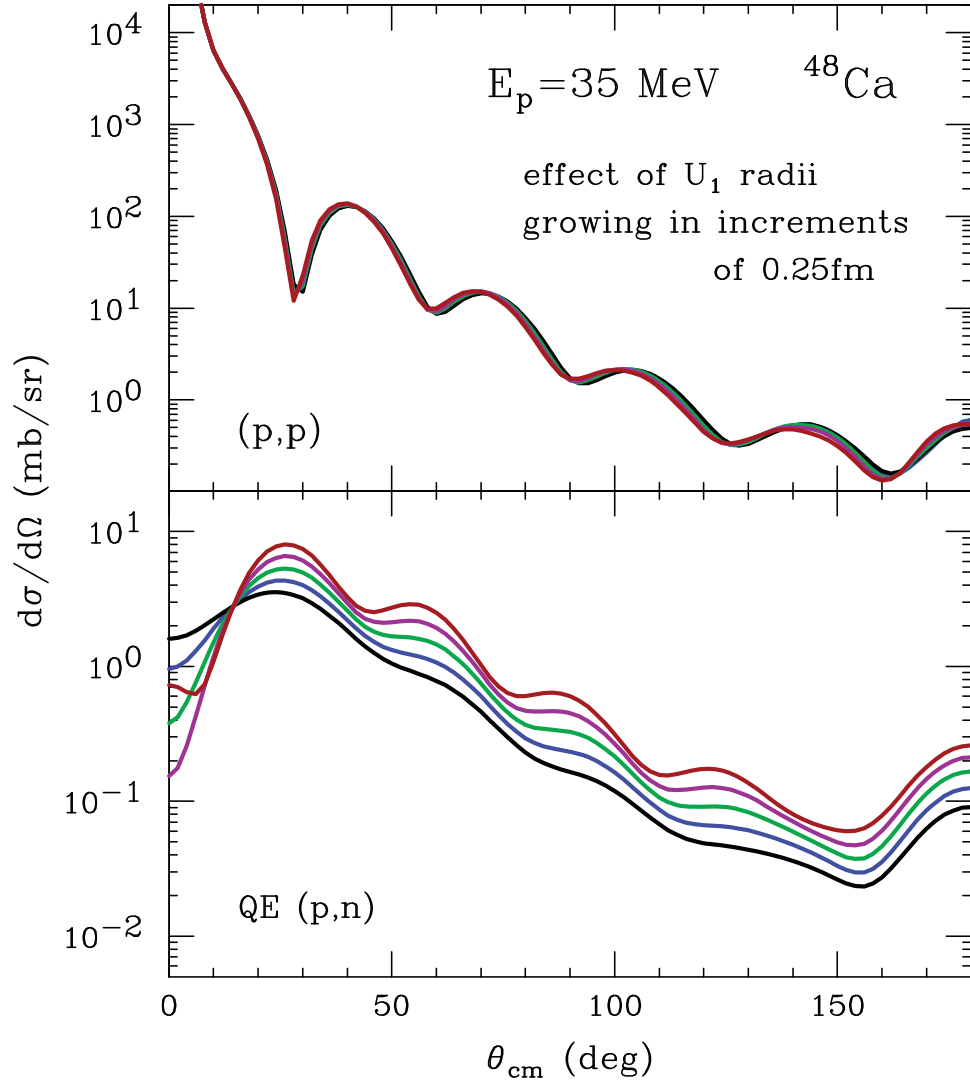
Robust prediction that in general

$$m_n^* > m_p^*$$

# Isvector optical potential and quasielastic charged-current reactions

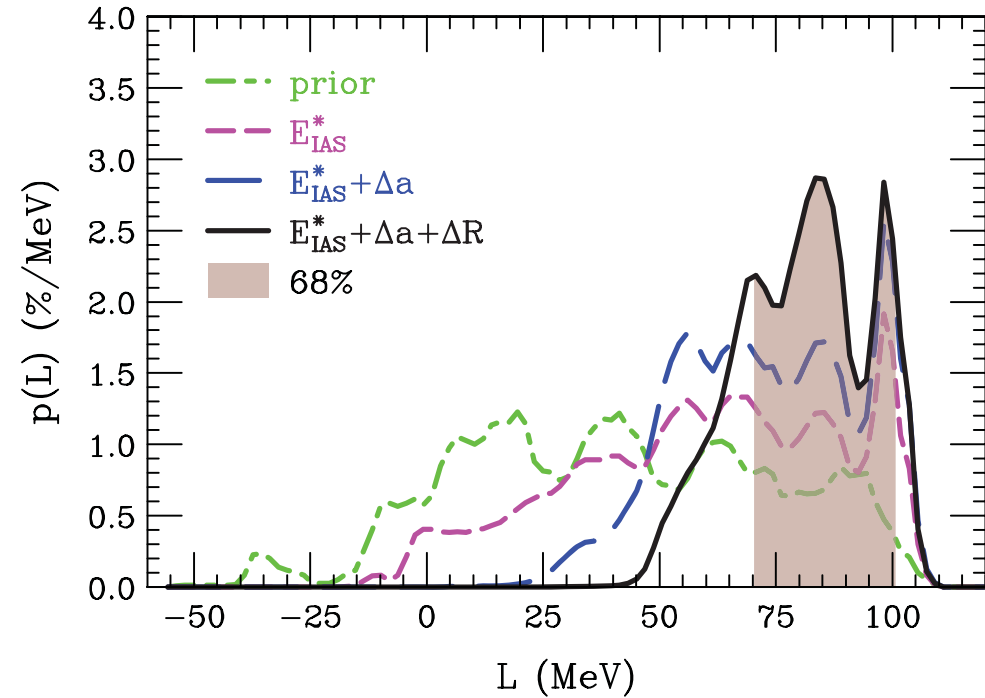


Danielewicz, Singh & Lee, NPA (2017)



- Charged-current reactions probe isovector part of optical potential

$$\langle n, Z + 1 | U(\mathbf{r}) | p, Z \rangle = 2 \frac{\sqrt{|N - Z|}}{A} U_1(\mathbf{r})$$



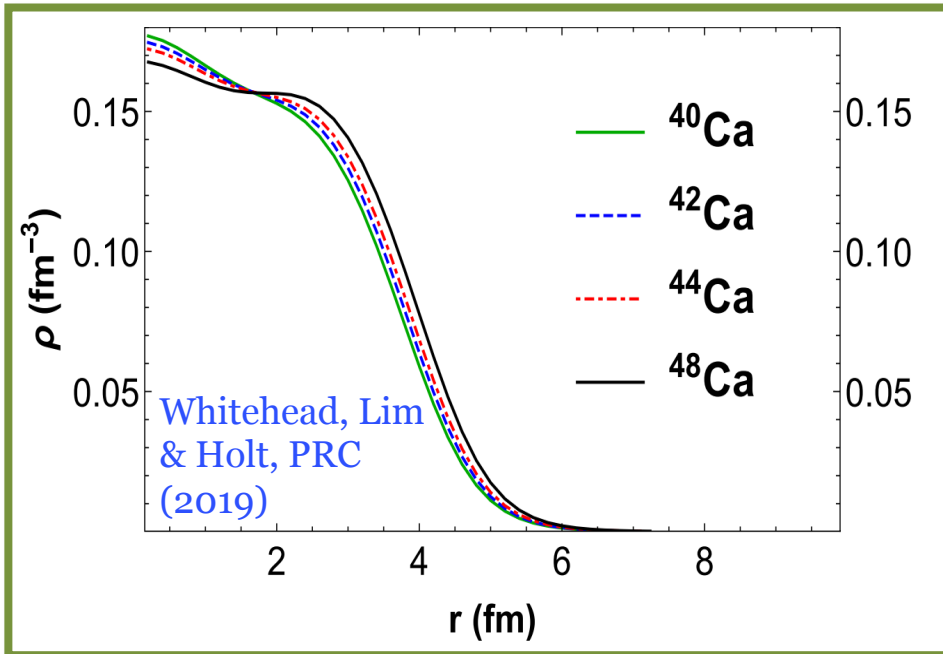
# From nuclear matter to *finite nuclei*: Local density approximation (LDA)

$$V(E; r) + iW(E; r) = V(E; k_f^p(r), k_f^n(r)) + iW(E; k_f^p(r), k_f^n(r))$$

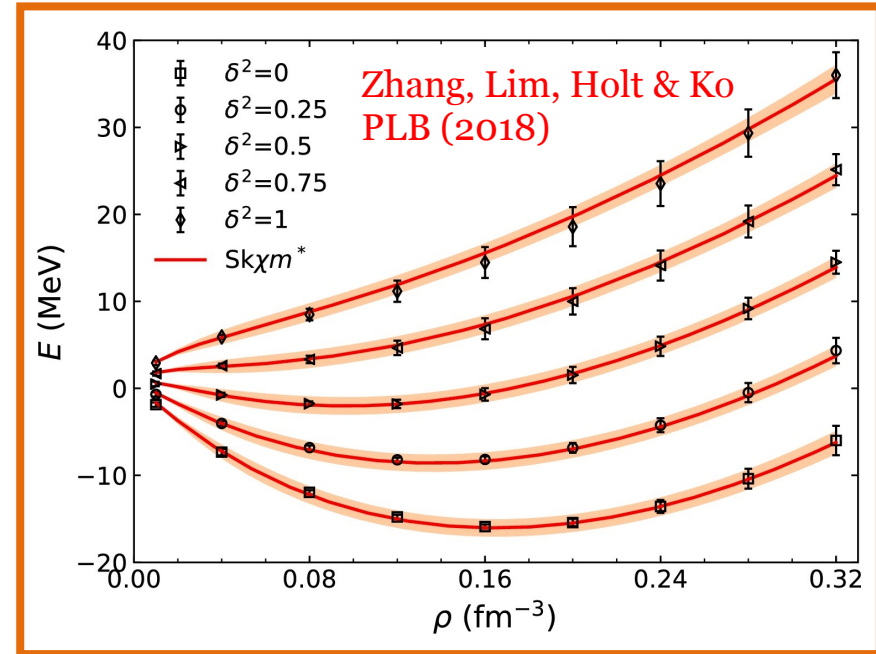
Local density approximation:

Optical potential in a finite nucleus matched to that of infinite matter at same isoscalar/isovector density

Density distributions from Sk $\chi$ 450 effective interaction



Sk $\chi$ 450 fitted to equation of state and effective mass

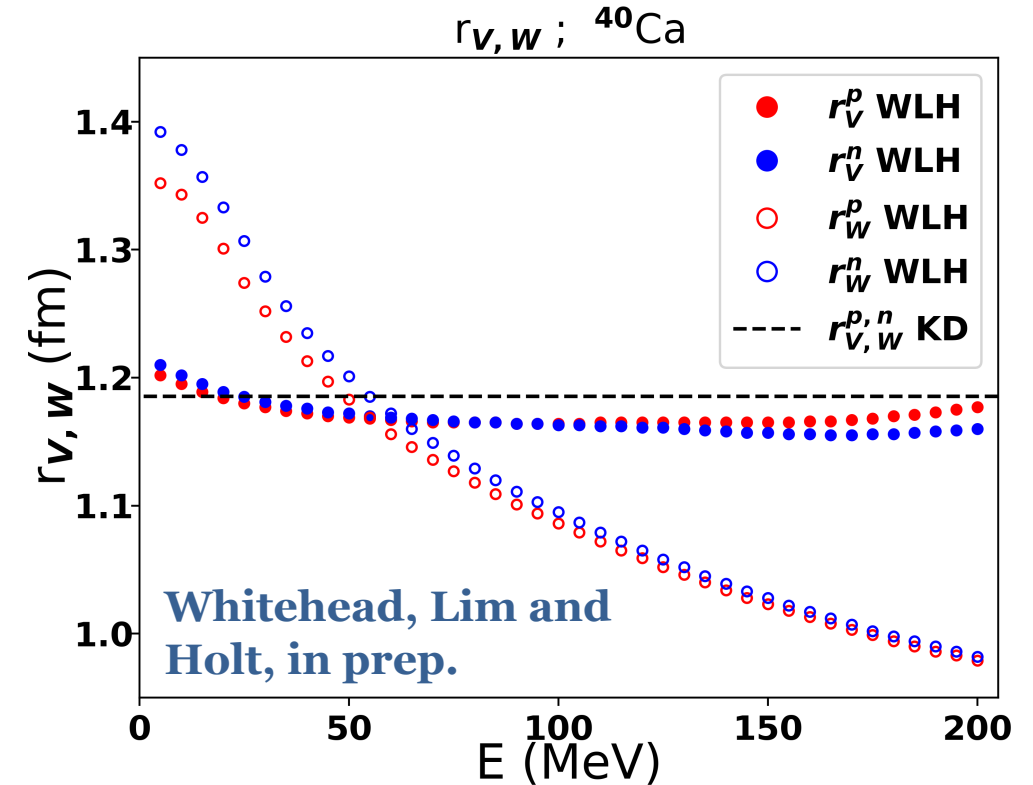


## New insights from microscopic calculations

- Phenomenological optical potentials make (necessary) assumptions in the parameterization
- Global analysis from microscopic calculations can **inform** phenomenological parameterizations
- Example: Energy-dependent Woods-Saxon shape parameters

$$\mathcal{V}(r, E) = V(E) \frac{1}{1 + e^{(r-r_0)/a}}$$

↑  
radius



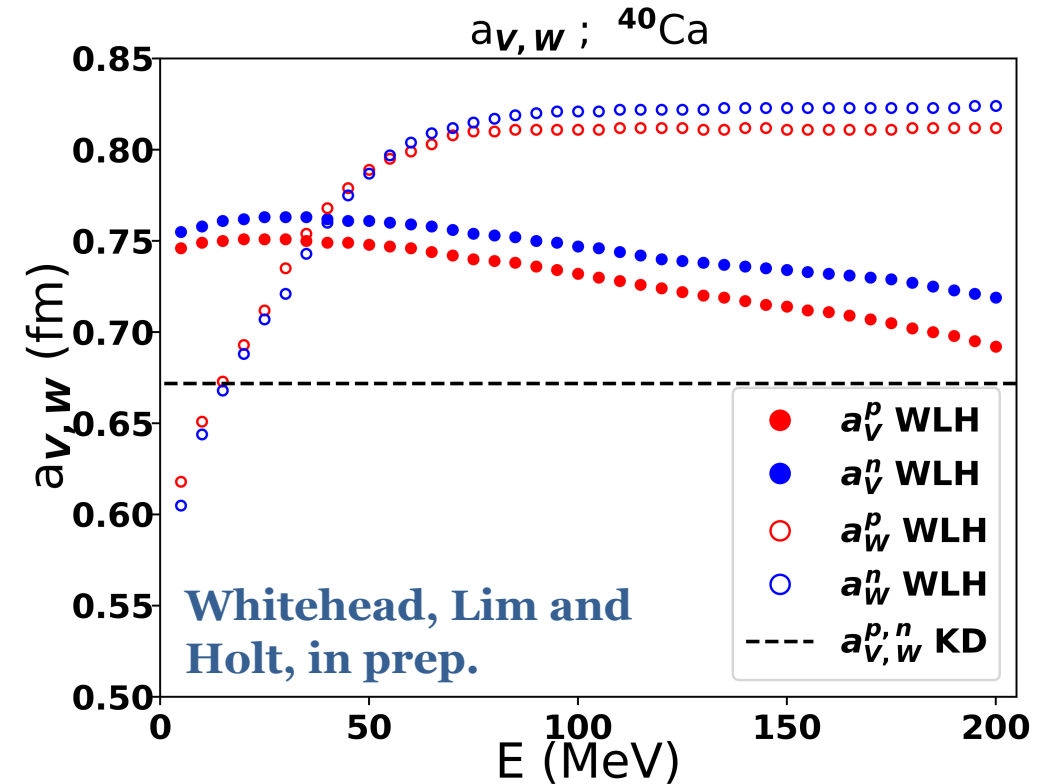


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$$\mathcal{V}(r, E) = V(E) \frac{1}{1 + e^{(r-r_0)/a}}$$

↑  
diffuseness



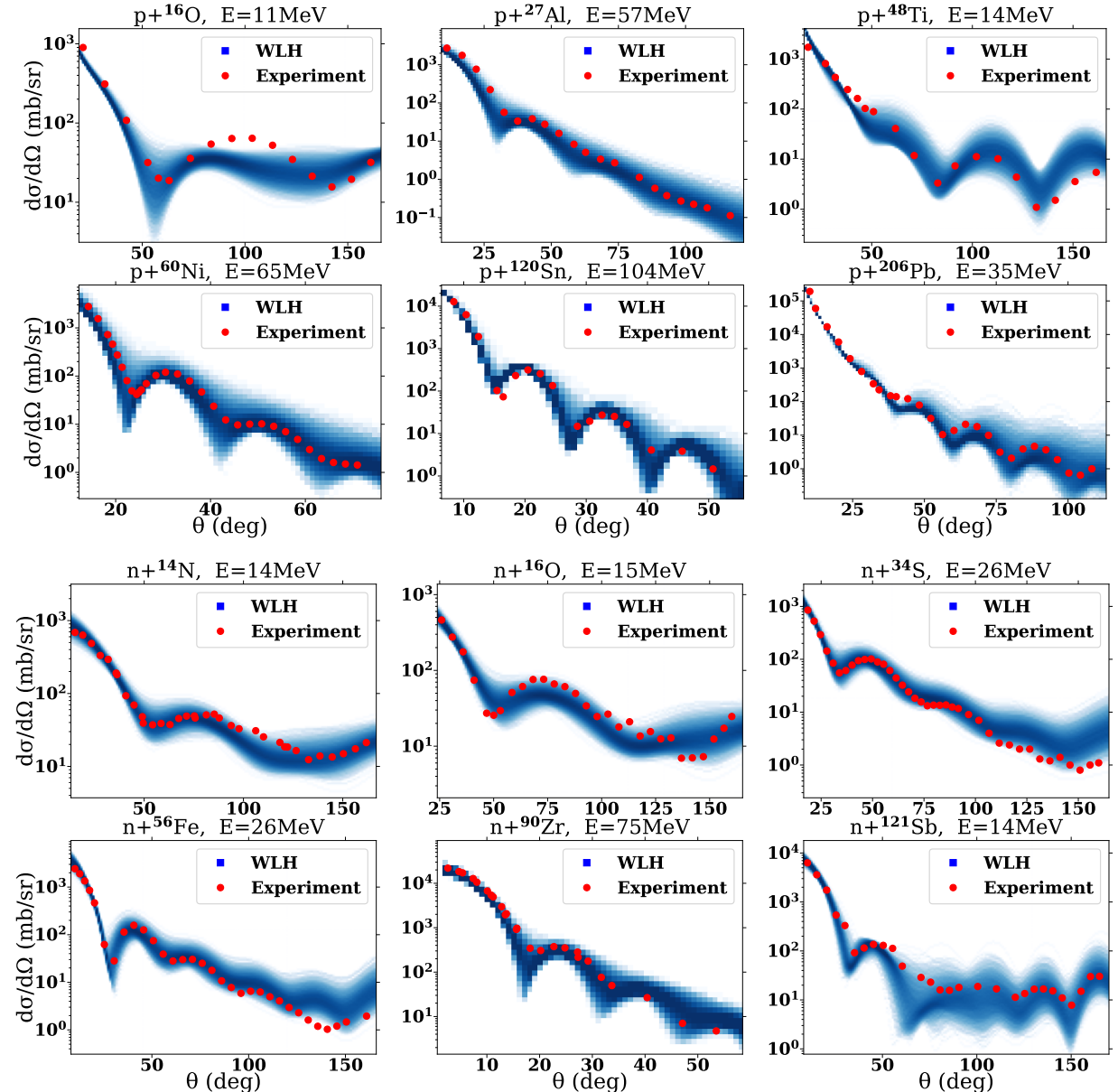
# Uncertainty quantification: differential elastic scattering cross sections



Whitehead, Lim & Holt, PRL (2021)

New global “WLH” microscopic global optical potential with uncertainties

- Proton/neutron optical potentials for **1800 target nuclei**
- Projectile energies  $E < 200$  MeV
- Microscopic results motivate new directions for phenomenology
  - Different Woods-Saxon geometry parameters for real and imaginary parts
  - Energy dependence of Woods-Saxon geometry parameters
- Uncertainties obtained using **5000 sampled global optical potentials** from covariance analysis of 5 chiral optical potentials



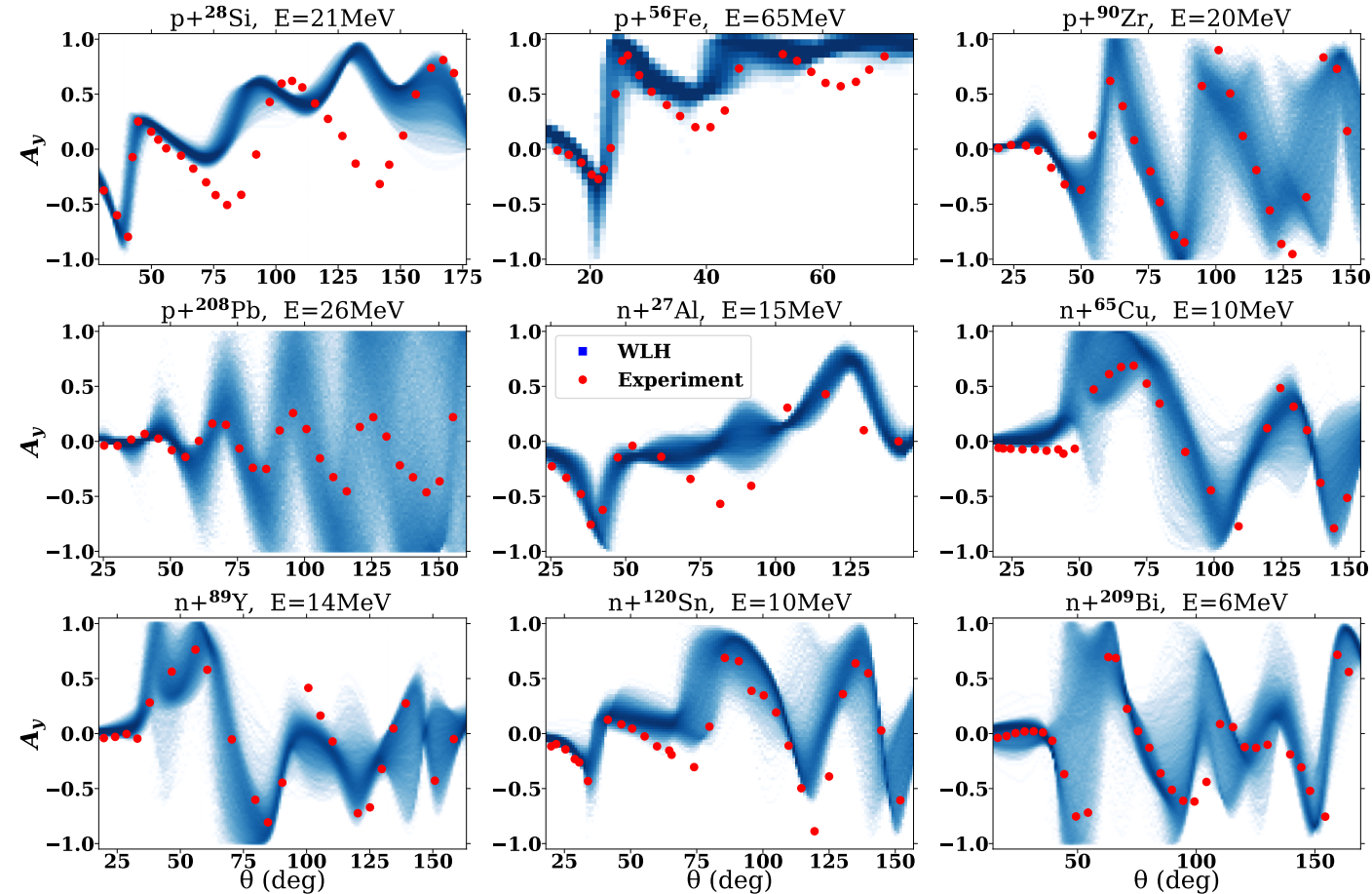
# Uncertainty quantification: vector analyzing powers



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Whitehead, Lim & Holt (2021)



Codes to generate optical potentials

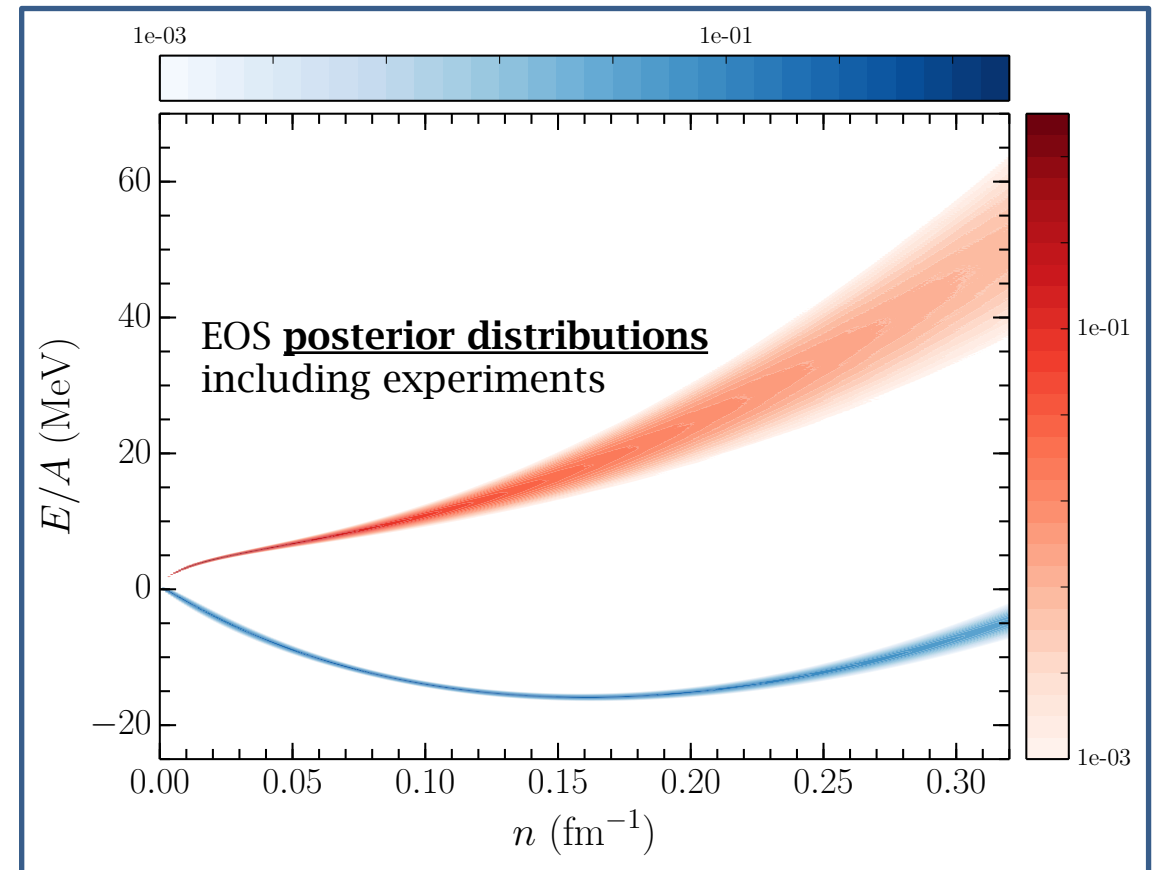
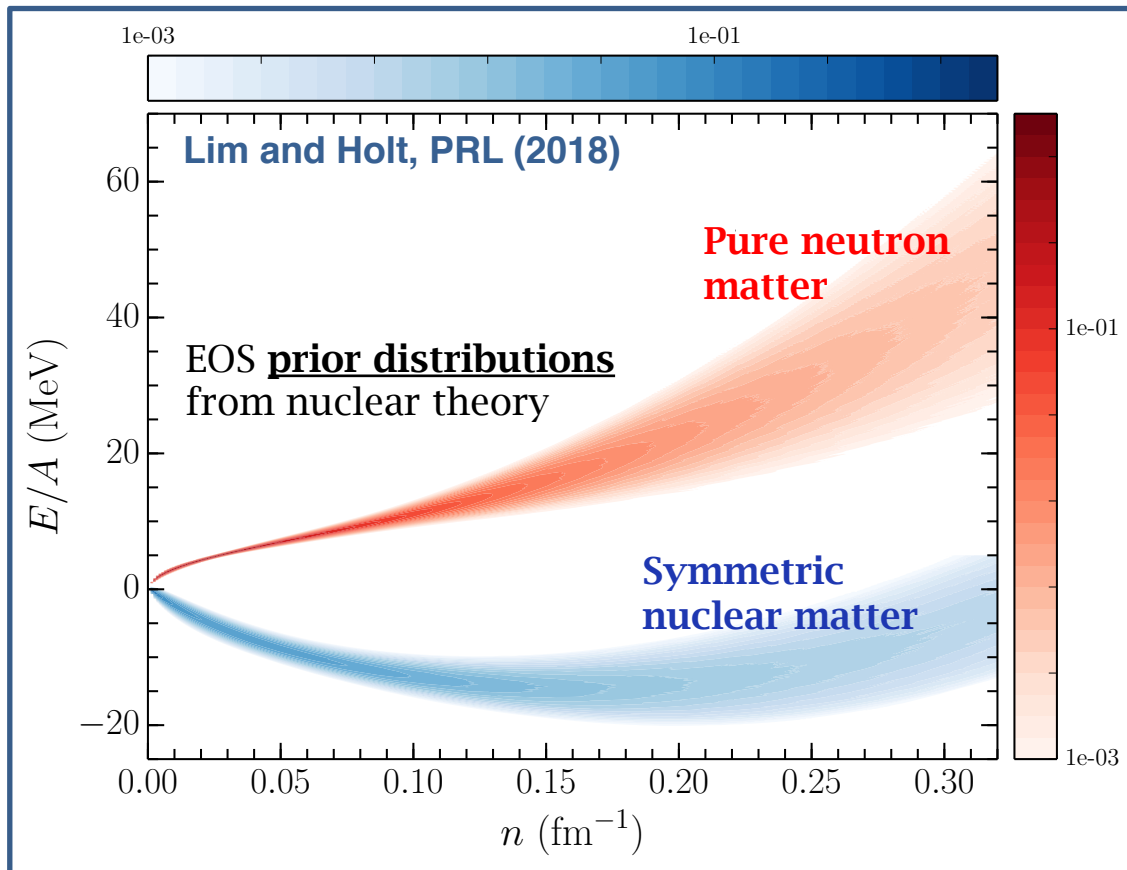
<https://www.trwhitehead.com/WLH>

# Consistency with the nuclear EOS



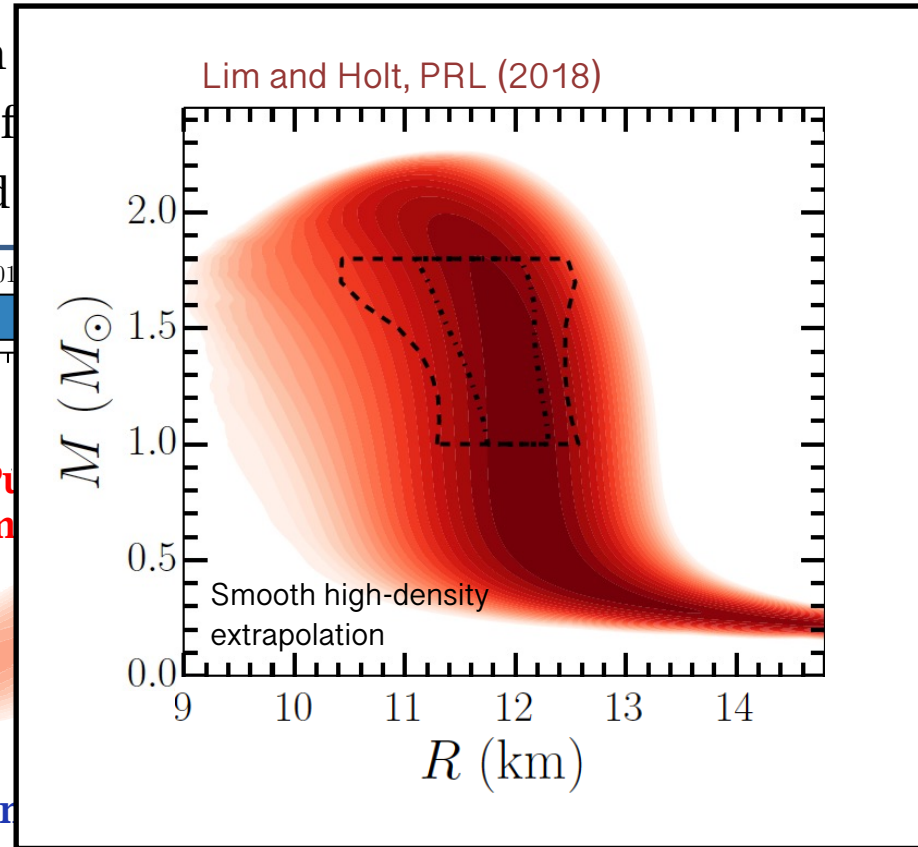
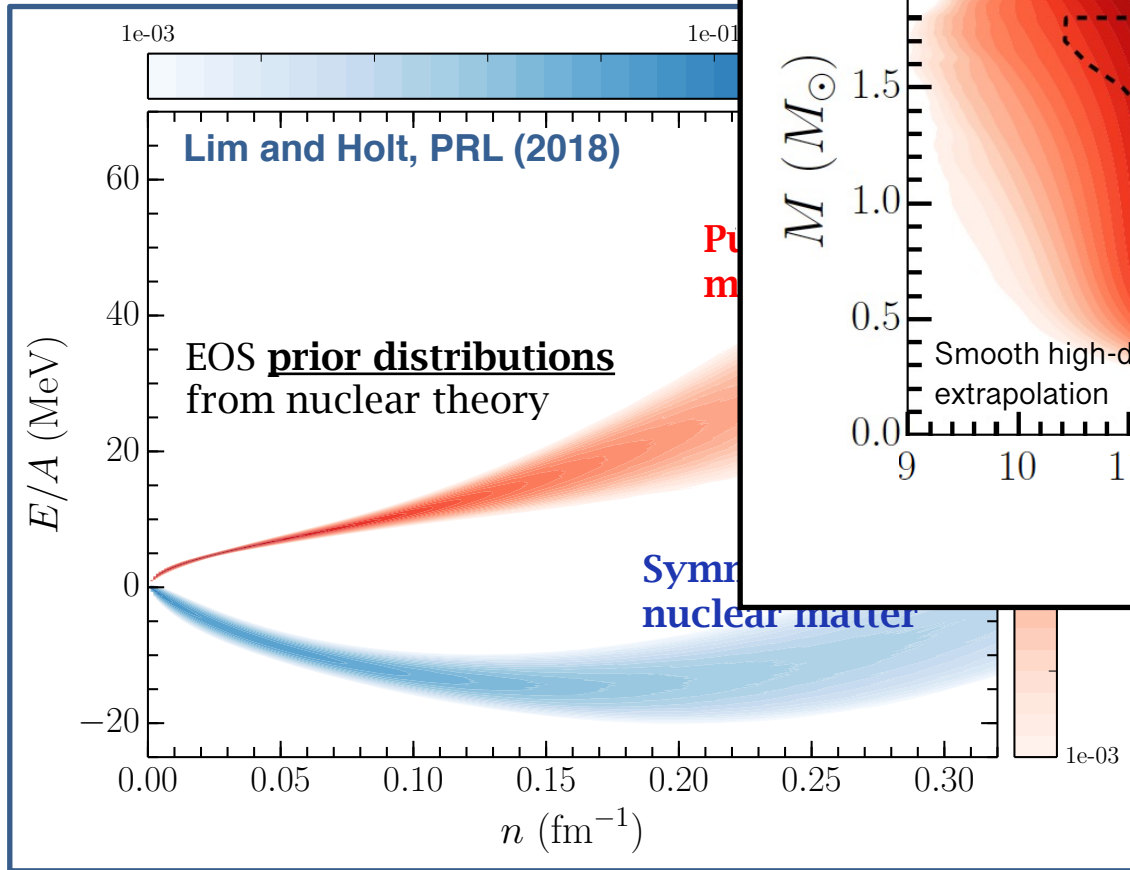
- Order in the chiral expansion
- Scale dependence of nuclear force
- Quantum many-body method

- Nuclear binding energies
- Experimental constraints on isospin-asymmetry energy

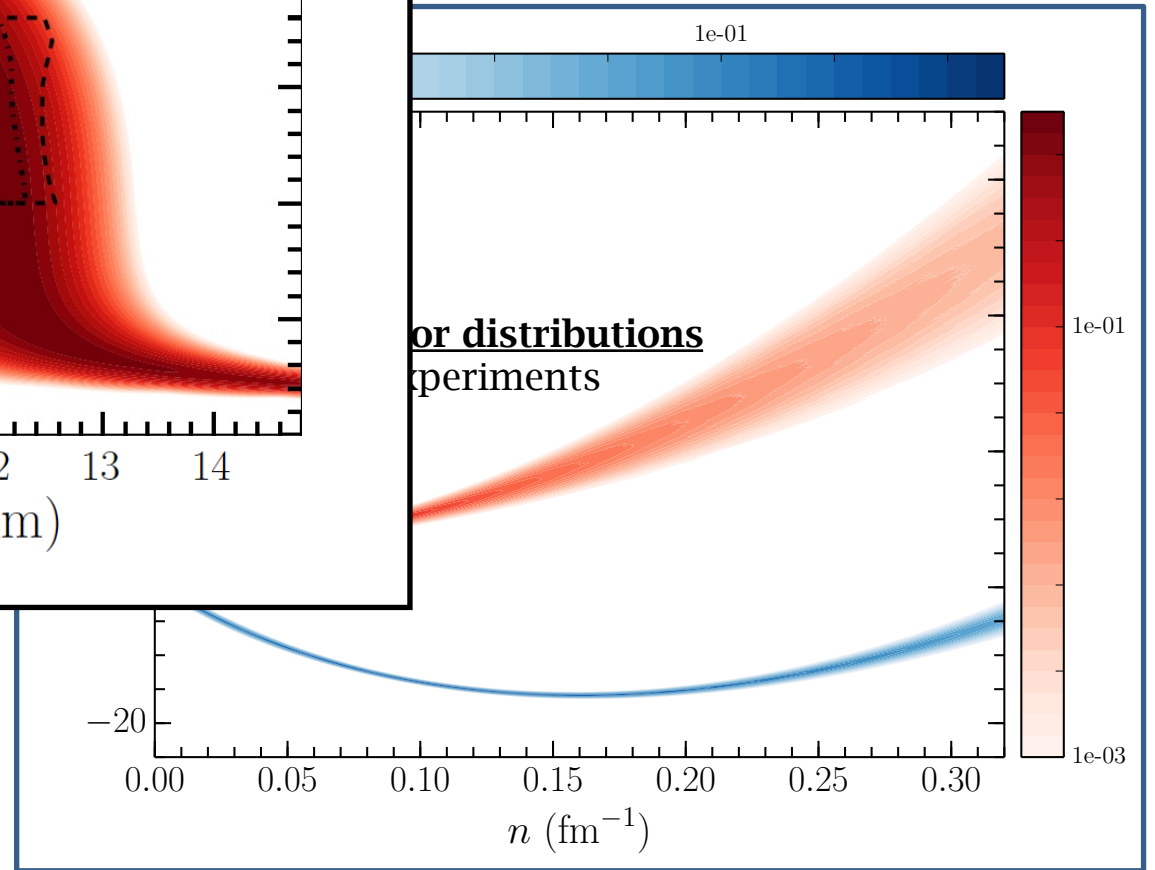


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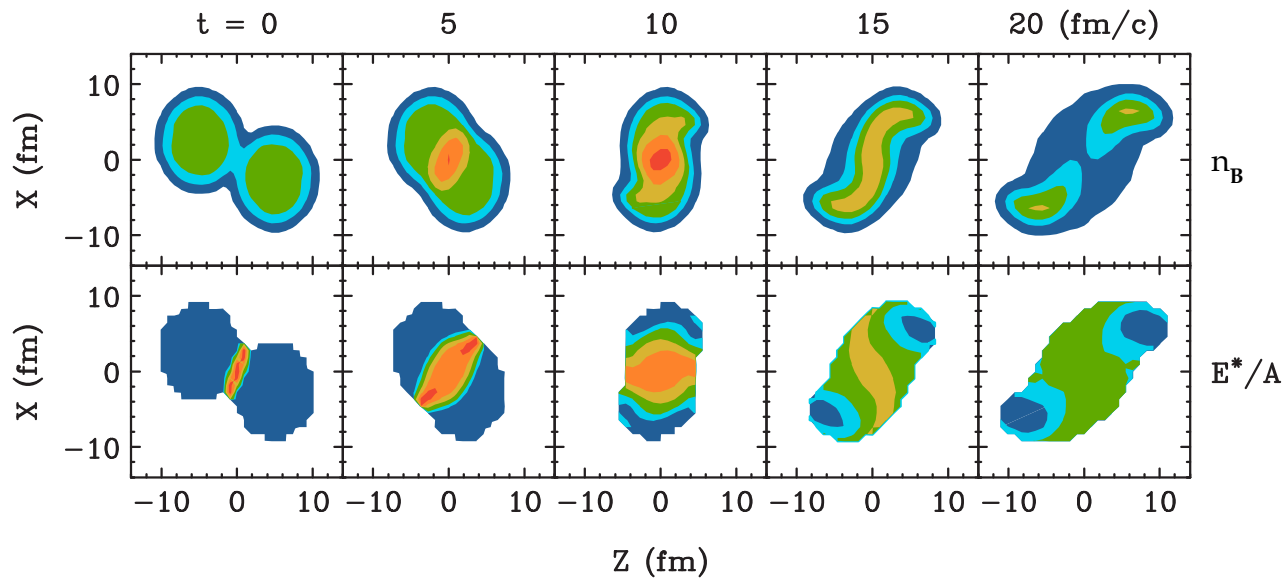
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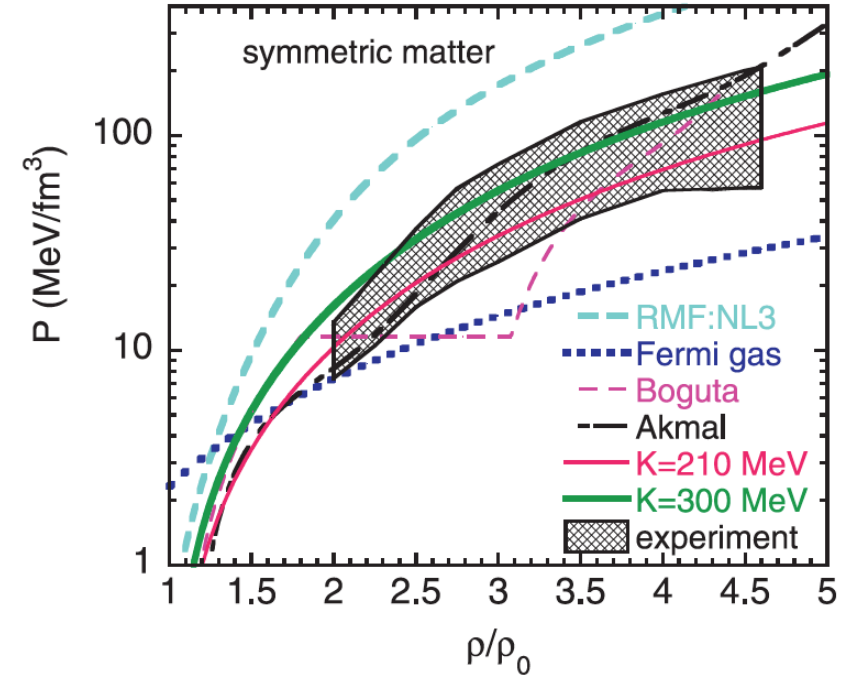
## Dense Nuclear Matter Equation of State from Heavy-Ion Collisions \*

arXiv:2301.13253

Agnieszka Sorensen<sup>1</sup>, Kshitij Agarwal<sup>2</sup>, Kyle W. Brown<sup>3,4</sup>, Zbigniew Chajecski<sup>5</sup>,  
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Danielewicz et al., Science (2002)



- Analysis requires molecular dynamics or transport models

$$\frac{\partial f}{\partial t} + \nabla_p \varepsilon \cdot \nabla_r f - \nabla_r \varepsilon \cdot \nabla_p f = I$$

- **Observables:** transverse and elliptic flow, fragment yields, charged pion ratios

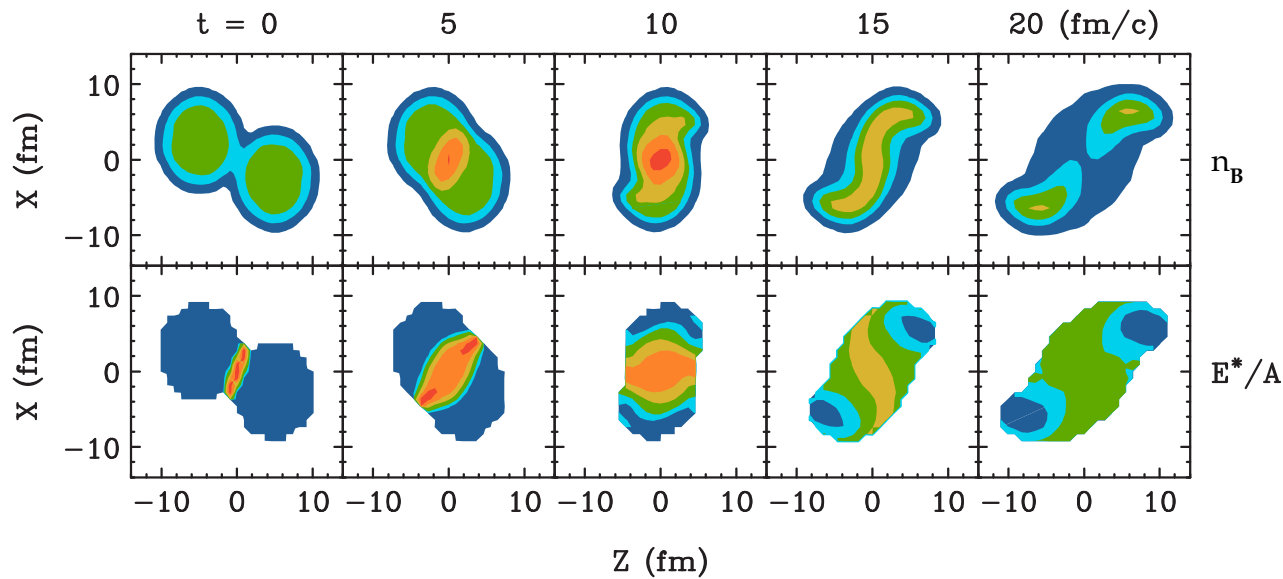
# Transport simulations of heavy-ion collisions and the EOS



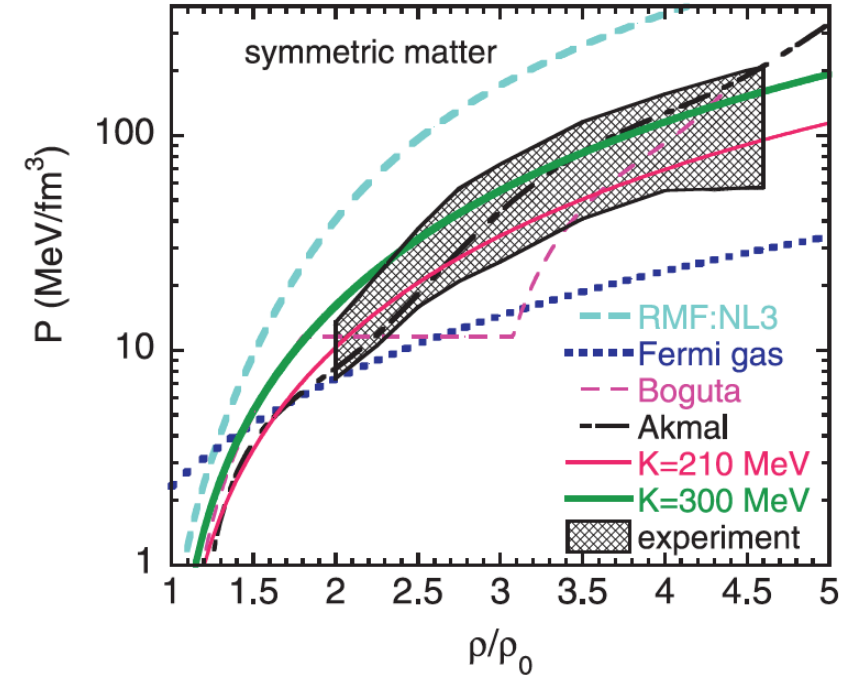
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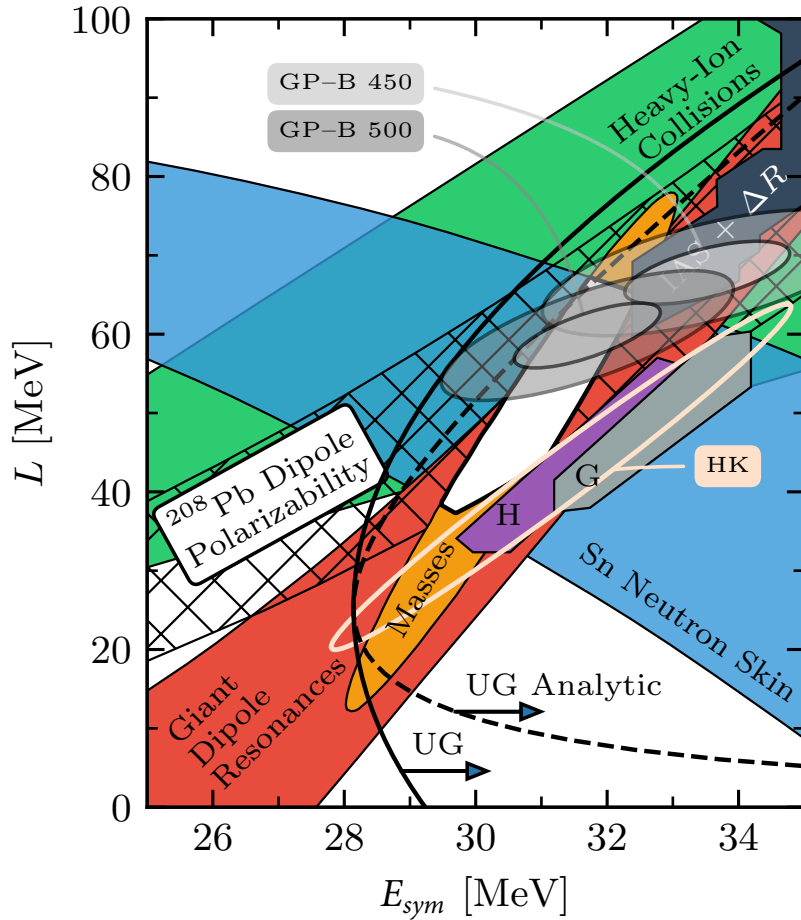
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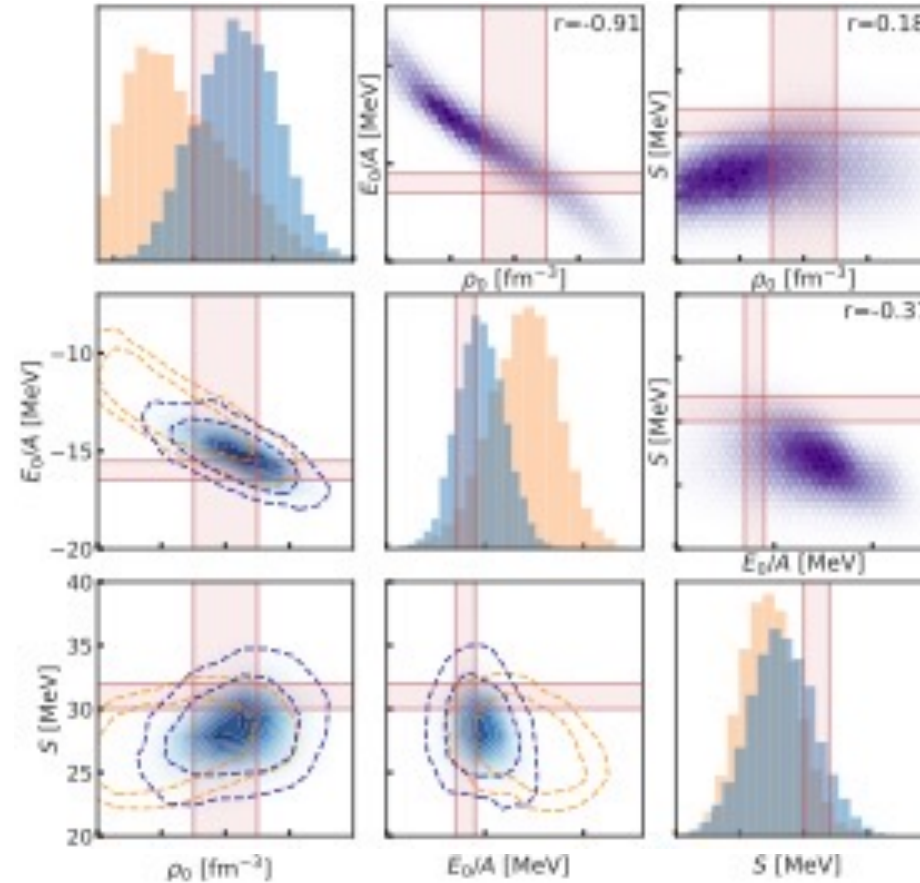
# Improved uncertainty quantification

Drischler, Furnstahl, Melendez & Phillips, PRL (2020)

Drischler, Holt & Wellenhofer, ARNPS (2021)



Jiang, Forssen, Djarv & Hagen, arXiv2212.13203



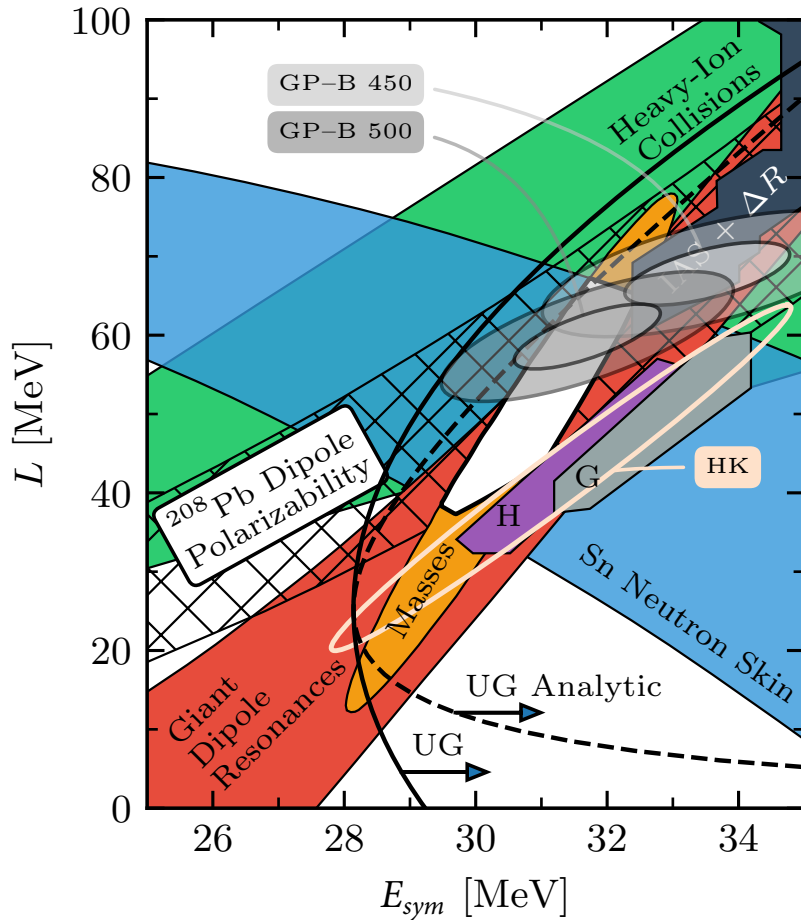
- EFT truncation errors
- Probability distributions for EFT low-energy constants



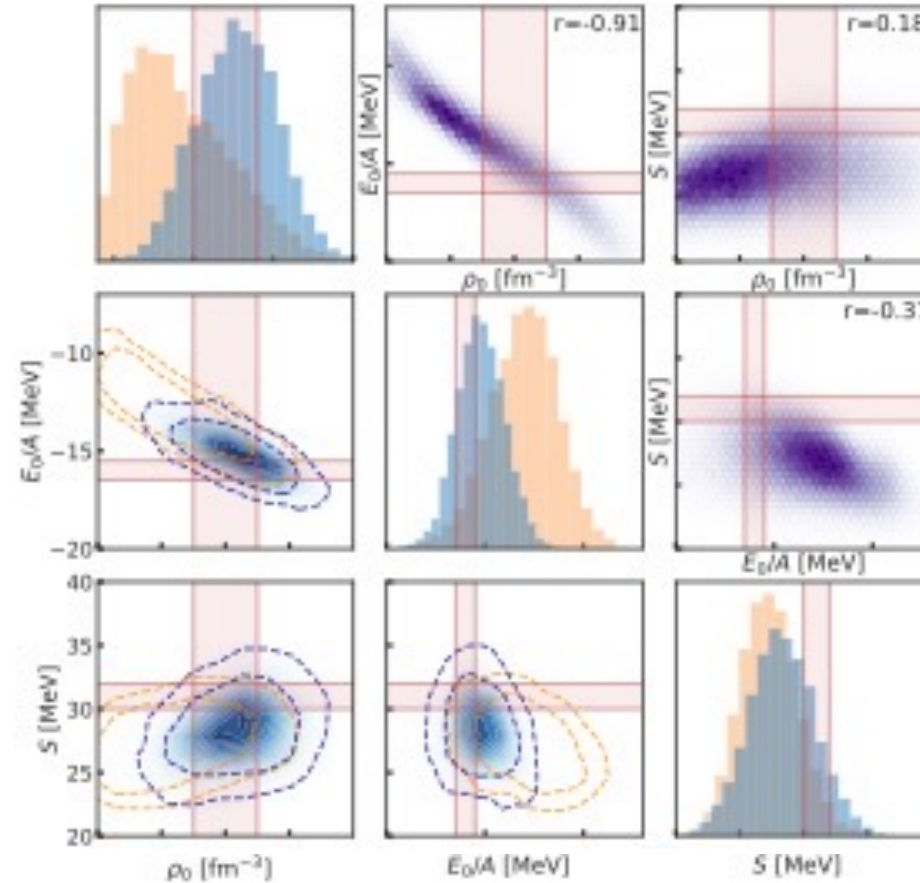
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**Goal: use generative machine learning models to create new chiral potentials**

# Generative modeling for nucleon-nucleon interactions



Glow: generate realistic faces

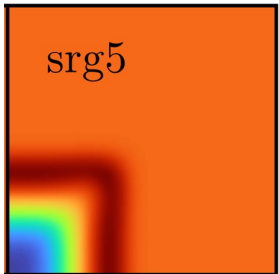
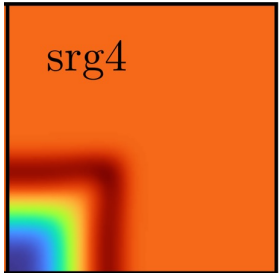
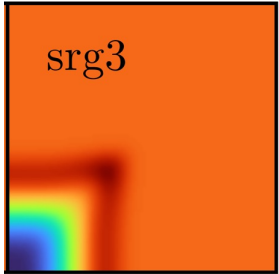
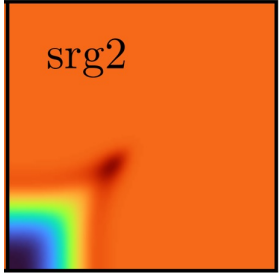


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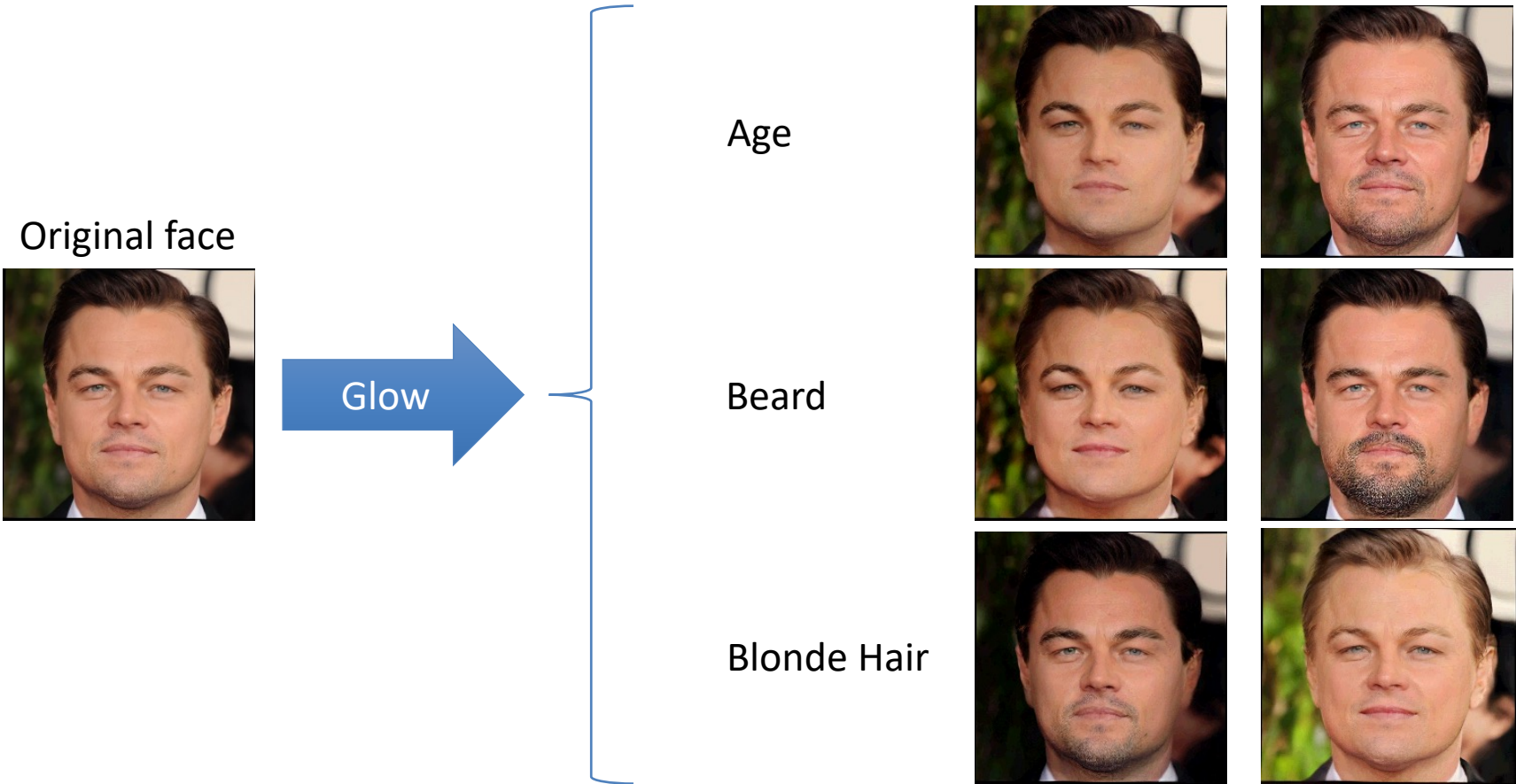


Glow: generate realistic faces

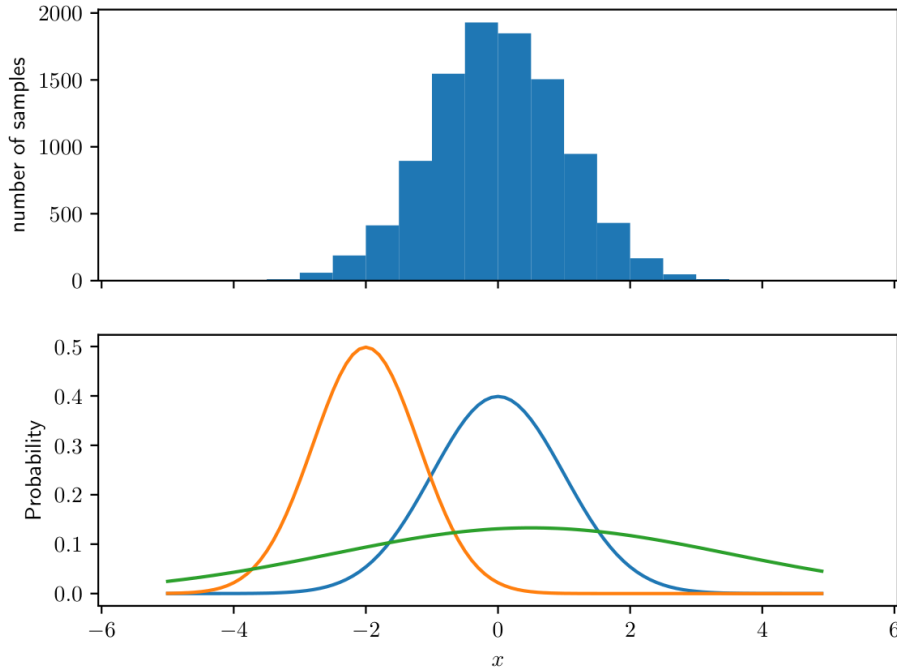
Nuclear Potentials



Glow: Manipulate attributes of a face ~~face~~ **Nuclear Potentials**



## Basic Algorithm



**Maximize likelihood:** optimize the estimation of density

Example:  $x \sim \mathcal{N}(\mu = 0, \sigma = 0.1)$

$$\mathcal{L}_1 = \prod_i \mathcal{N}(X_i | \mu = 0, \sigma = 0.1)$$

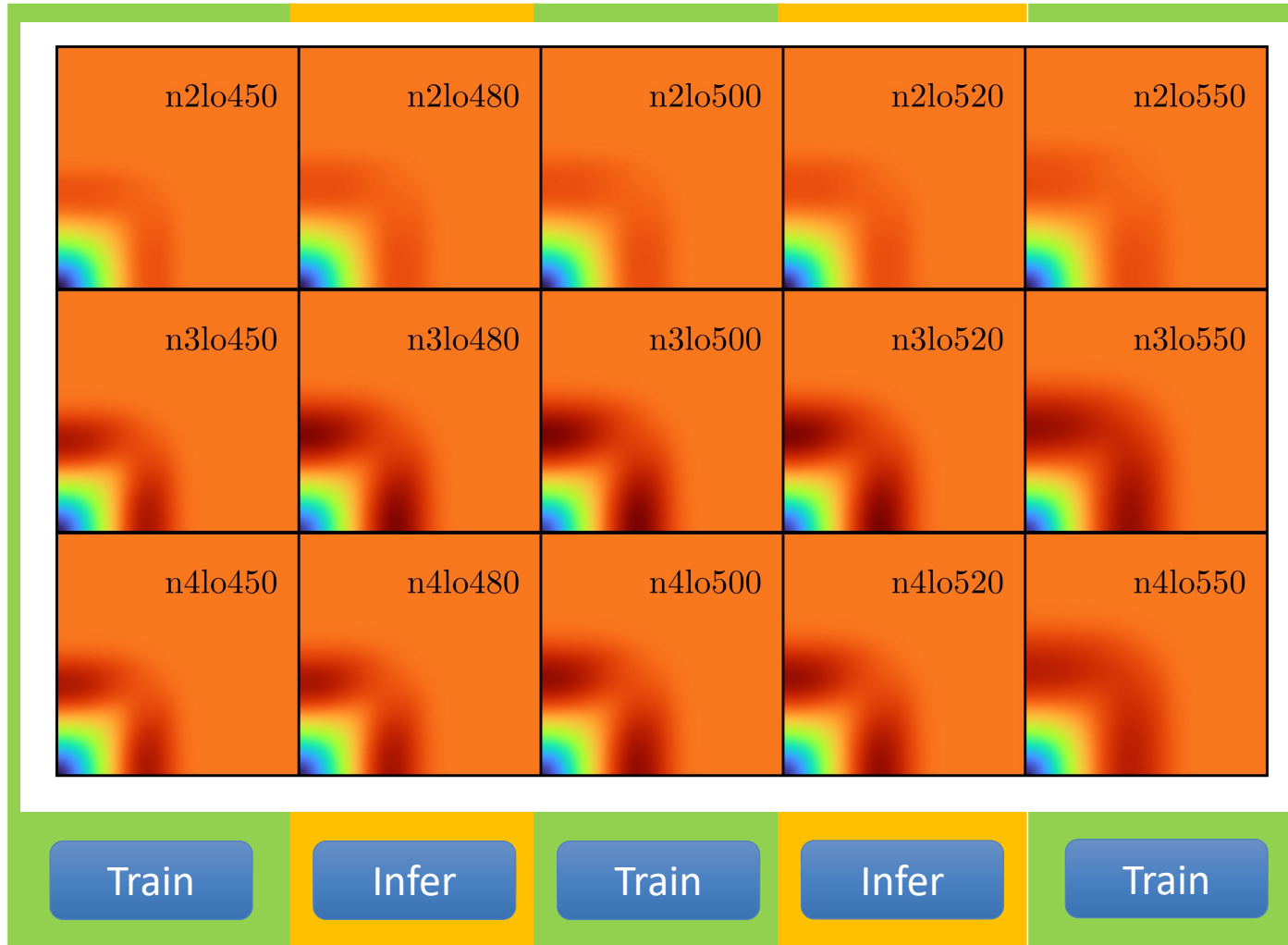
$$\mathcal{L}_2 = \prod_i \mathcal{N}(X_i | \mu = -2, \sigma = 0.8)$$

$$\mathcal{L}_3 = \prod_i \mathcal{N}(X_i | \mu = 0.5, \sigma = 3)$$

A proper probability distribution model  $\mathcal{N}(\mu, \sigma) \rightarrow$  the maximum likelihood  $\mathcal{L}$

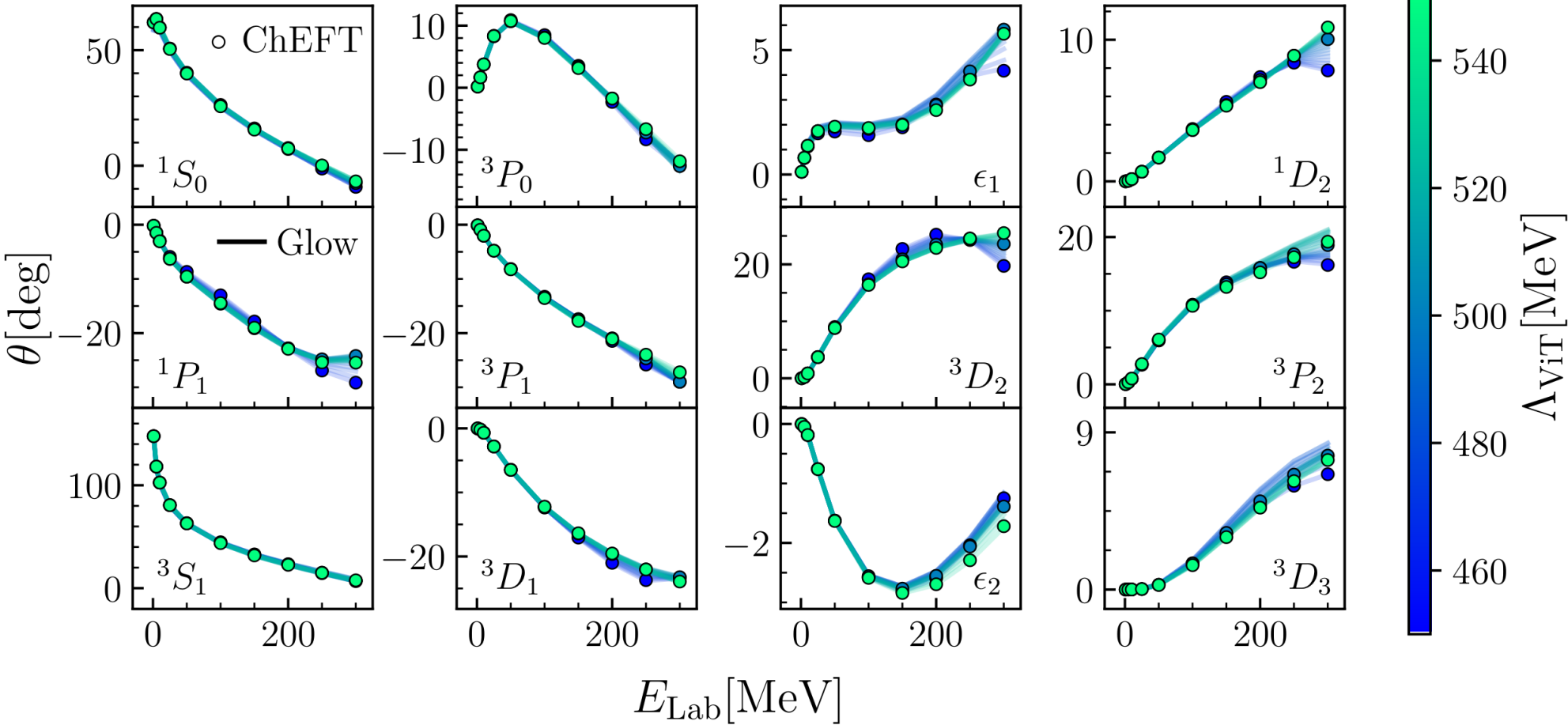
If  $X =$  potential  $V$ , how to find  $p(V) \rightarrow$  maximum  $\mathcal{L}(V)$ ?

# Generative modeling for nucleon-nucleon interactions



1. Distinguish truncation orders in ChEFT
2. Predict how the chiral potentials evolve with cutoff

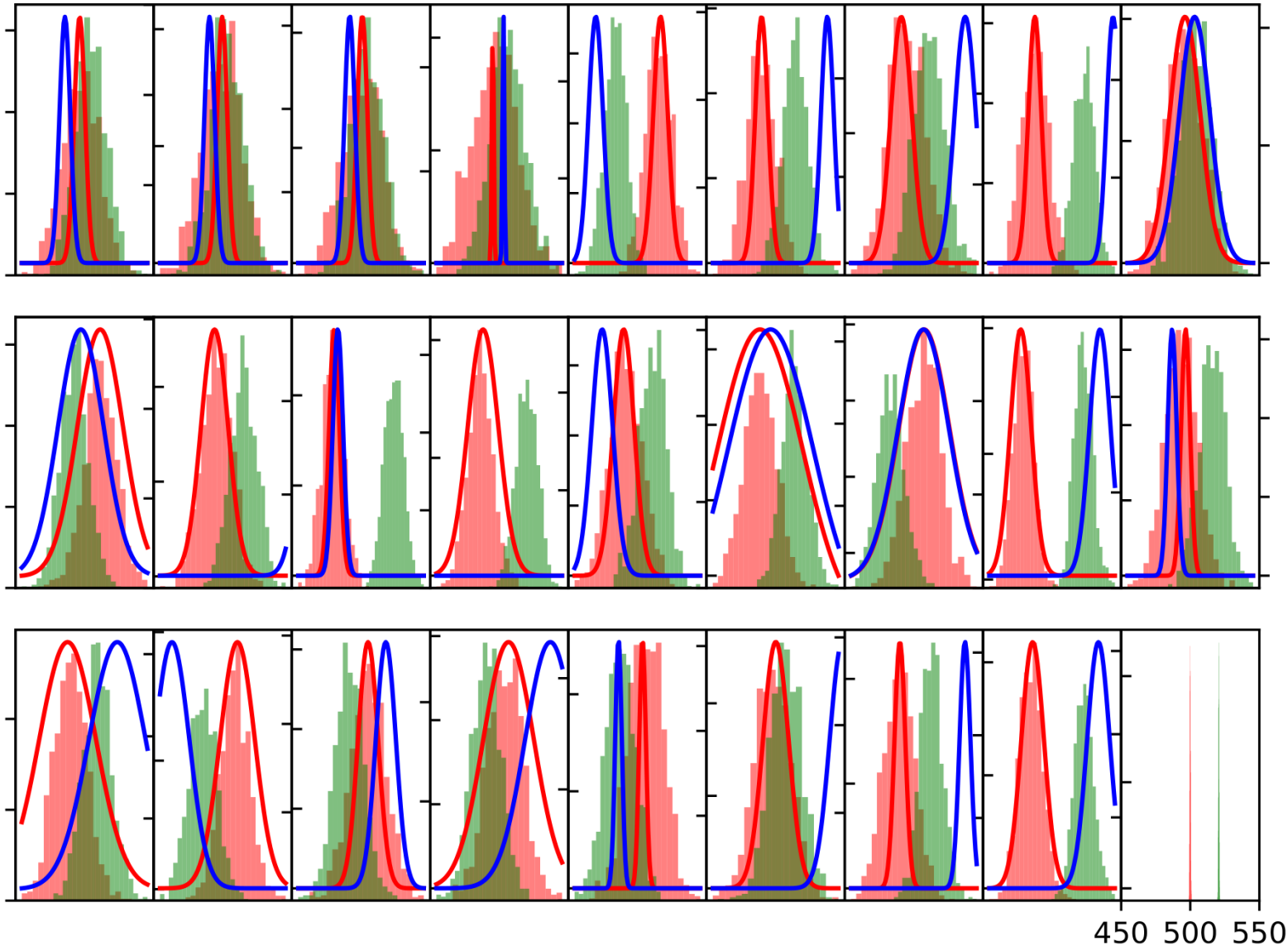
# Generative modeling for nucleon-nucleon interactions



# Generative modeling for nucleon-nucleon interactions



Low energy constant distributions



1. Can extract LECs from generated chiral potentials
2. Input LEC distributions can also be propagated to

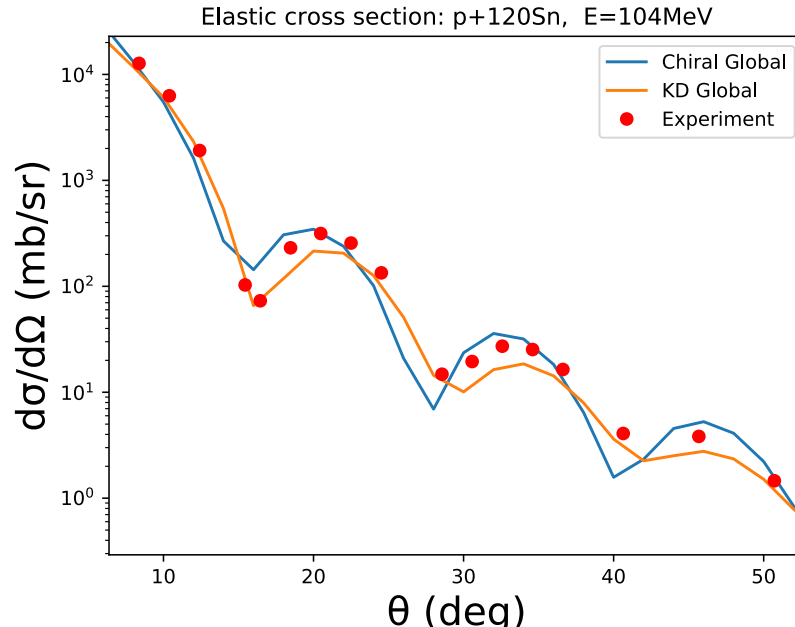
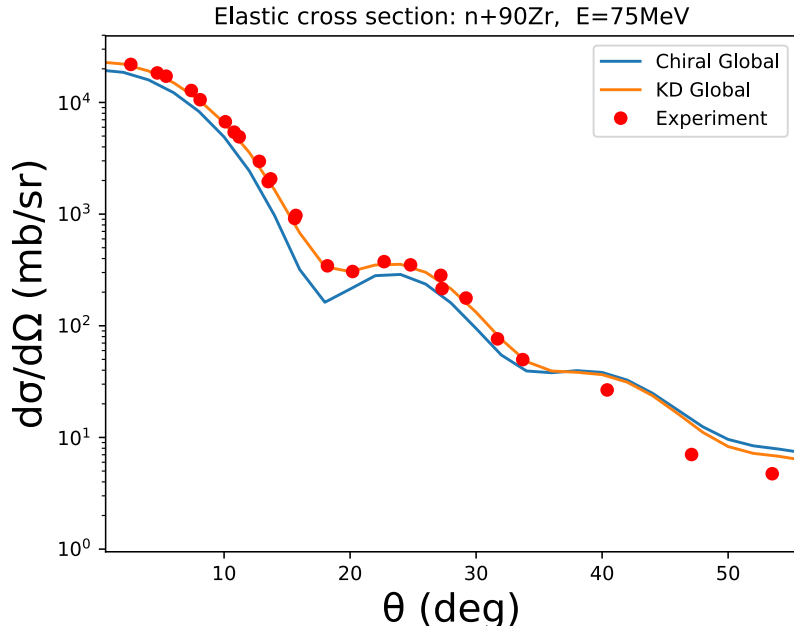
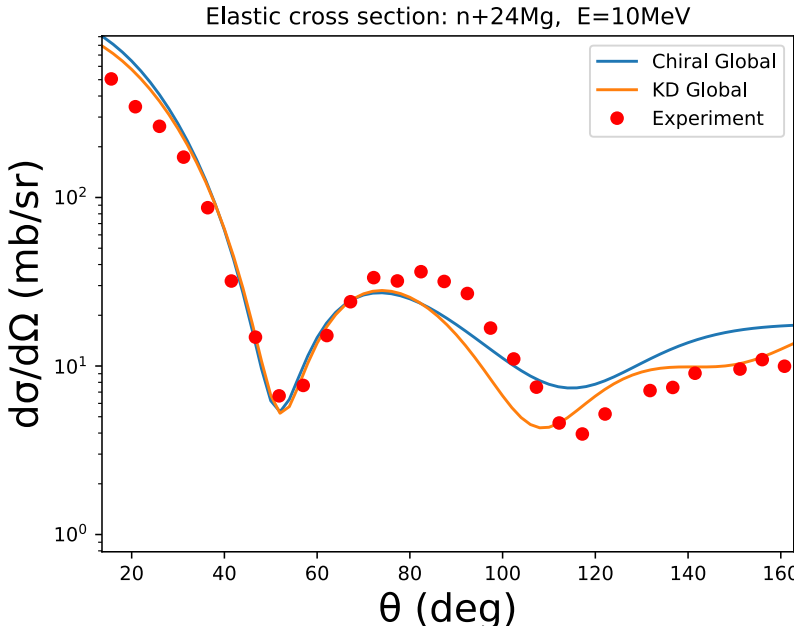


# Summary and future directions

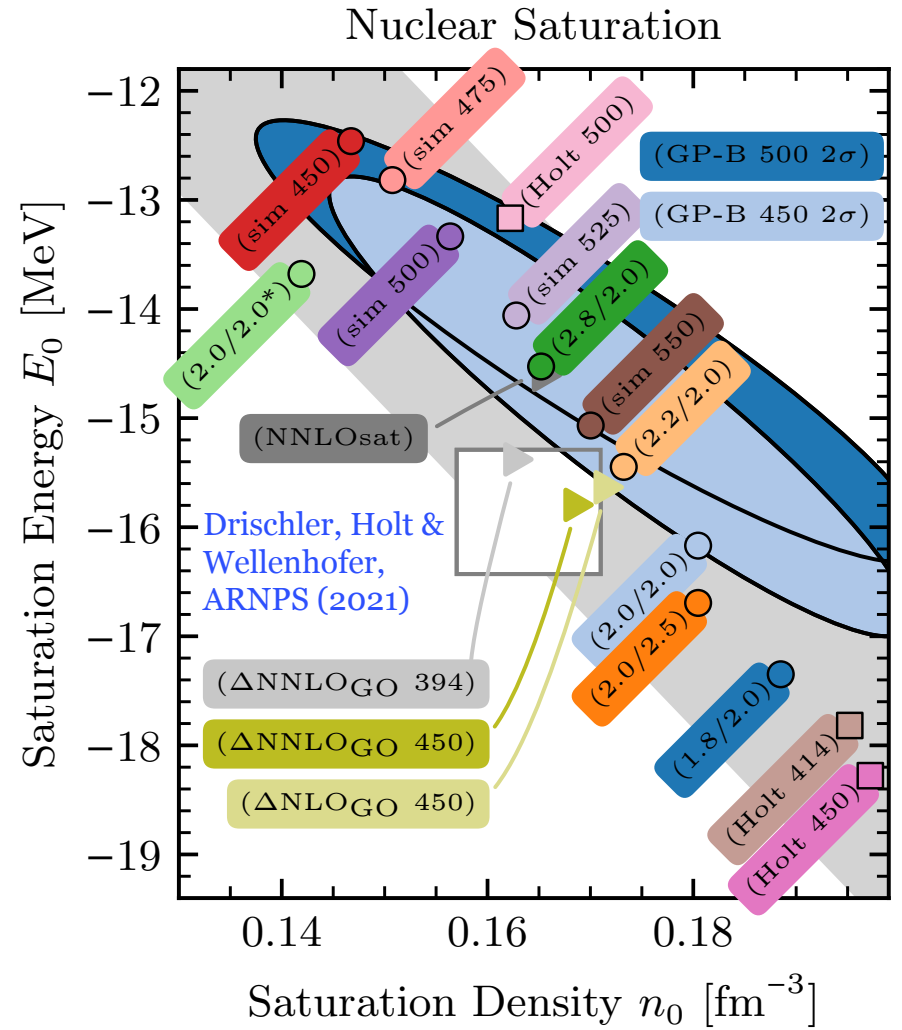
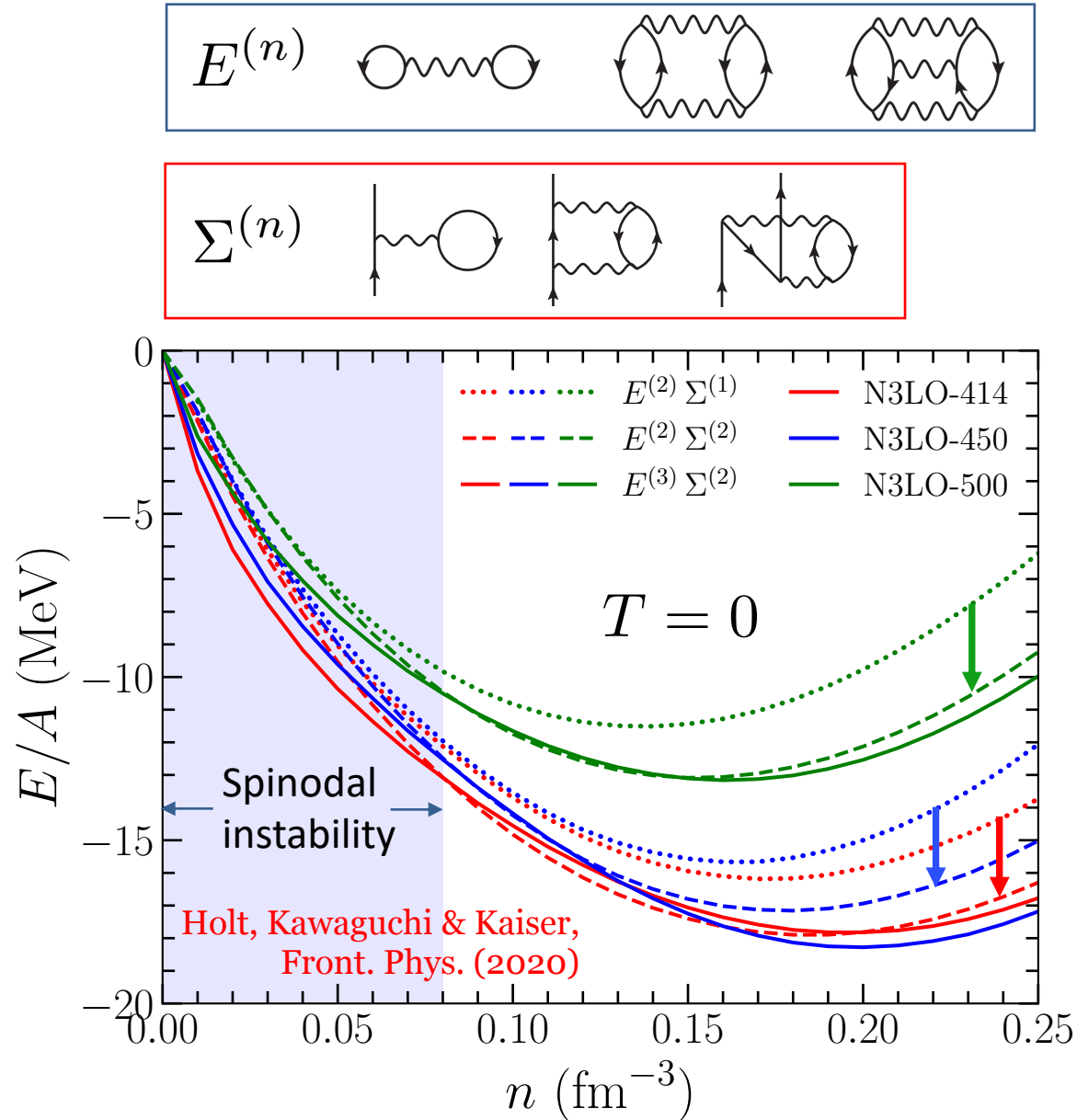


- First microscopic global optical potential with quantified uncertainties
  - Good description of differential nucleon-nucleus elastic scattering cross sections within uncertainties
  - Analyzing powers have larger uncertainties but also compare well to experiment
- Improved uncertainty quantification possible: variations in chiral low-energy constants, more sophisticated treatments of EFT truncation errors and choice of resolution scale
- Work in progress to interface new optical potentials with nuclear reaction codes for rare-isotope beam experiments

# Comparison to phenomenology and data: selected results



# Nuclear matter uncertainties



● Saturation point is a fine-tuned quantity

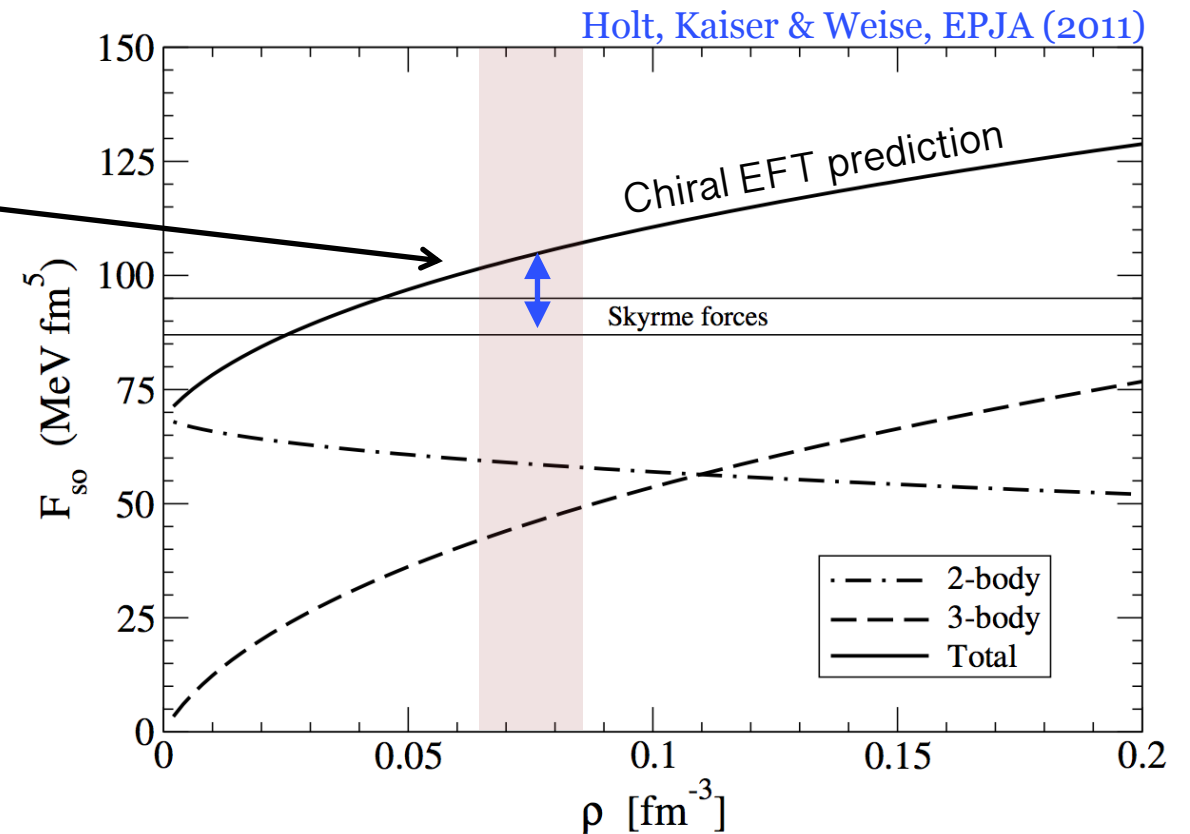
# From nuclear matter to *finite nuclei*: spin-orbit interactions

- Spin orbit interaction vanishes in infinite homogeneous nuclear matter

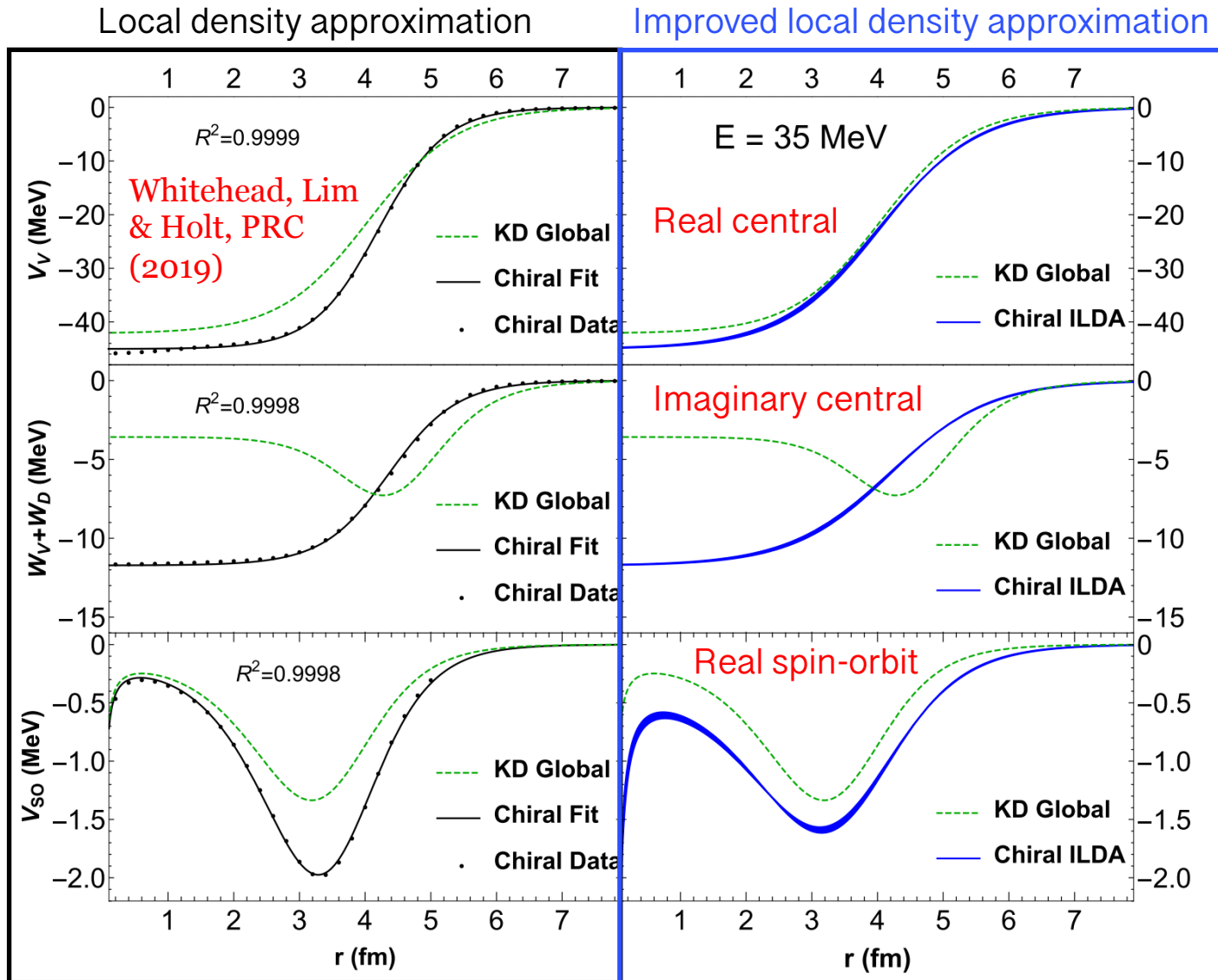
Density matrix expansion (Negele & Vautherin, PRC 1972)

$$E_{\text{HF}} = \frac{1}{2} \text{Tr}_1 \text{Tr}_2 \int d\vec{r}_1 \cdots d\vec{r}_4 \langle \vec{r}_1 \vec{r}_2 | V(1 - P_{12}) | \vec{r}_3 \vec{r}_4 \rangle \rho(\vec{r}_3, \vec{r}_1) \rho(\vec{r}_4, \vec{r}_2)$$

- 20% too large at density region of interest
- 2<sup>nd</sup>-order perturbative contributions known to reduce spin-orbit strength



# From nuclear matter to *finite nuclei*: Improved LDA



$$V(E; r)_{ILDA} = \frac{1}{(t\sqrt{\pi})^3} \int V(E; r') e^{-\frac{|\vec{r}-\vec{r}'|^2}{t^2}} d^3 r'$$

- Finite range of nuclear force must be accounted for
- Introduce Gaussian smearing function with range parameter  $t$
- Increases the Woods-Saxon diffuseness parameter in agreement with phenomenology