

MIDTERM EXAM,

PHYSICS 852, Spring 2022

Friday, March 4, 11:30-12:20 PM

Your Name: _____

$$\int_{-\infty}^{\infty} dx e^{-x^2/(2a^2)} = a\sqrt{2\pi},$$

$$H = i\hbar\partial_t, \vec{P} = -i\hbar\nabla,$$

$$\sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix},$$

$$U(t, -\infty) = 1 + \frac{-i}{\hbar} \int_{-\infty}^t dt' V(t')U(t', -\infty),$$

$$\langle x|x'\rangle = \delta(x - x'), \langle p|p'\rangle = \frac{1}{2\pi\hbar}\delta(p - p'),$$

$$|p\rangle = \int dx |x\rangle e^{ipx/\hbar}, |x\rangle = \int \frac{dp}{2\pi\hbar} |p\rangle e^{-ipx/\hbar},$$

$$H = \frac{P^2}{2m} + \frac{1}{2}m\omega^2 x^2 = \hbar\omega(a^\dagger a + 1/2),$$

$$a^\dagger = \sqrt{\frac{m\omega}{2\hbar}}X - i\sqrt{\frac{1}{2\hbar m\omega}}P,$$

$$\psi_0(x) = \frac{1}{(\pi b^2)^{1/4}} e^{-x^2/2b^2}, b^2 = \frac{\hbar}{m\omega},$$

$$\rho(\vec{r}, t) = \psi^*(\vec{r}_1, t_1)\psi(\vec{r}_2, t_2)$$

$$\vec{j}(\vec{r}, t) = \frac{-i\hbar}{2m}(\psi^*(\vec{r}, t)\nabla\psi(\vec{r}, t) - (\nabla\psi^*(\vec{r}, t))\psi(\vec{r}, t)) - \frac{e\vec{A}}{mc}|\psi(\vec{r}, t)|^2.$$

$$H = \frac{(\vec{P} - e\vec{A}/c)^2}{2m} + e\Phi,$$

For $V = \beta\delta(x - y)$: $-\frac{\hbar^2}{2m}(\partial_x\psi(x)|_{y+\epsilon} - \partial_x\psi(x)|_{y-\epsilon}) = -\beta\psi(y),$

$$\vec{E} = -\nabla\Phi - \frac{1}{c}\partial_t\vec{A}, \vec{B} = \nabla \times \vec{A},$$

$$\omega_{\text{cyclotron}} = \frac{eB}{mc},$$

$$e^{A+B} = e^A e^B e^{-C/2}, \text{ if } [A, B] = C, \text{ and } [C, A] = [C, B] = 0,$$

$$Y_{0,0} = \frac{1}{\sqrt{4\pi}}, Y_{1,0} = \sqrt{\frac{3}{4\pi}}\cos\theta, Y_{1,\pm 1} = \mp\sqrt{\frac{3}{8\pi}}\sin\theta e^{\pm i\phi},$$

$$Y_{2,0} = \sqrt{\frac{5}{16\pi}}(3\cos^2\theta - 1), Y_{2,\pm 1} = \mp\sqrt{\frac{15}{8\pi}}\sin\theta\cos\theta e^{\pm i\phi},$$

$$Y_{2,\pm 2} = \sqrt{\frac{15}{32\pi}}\sin^2\theta e^{\pm 2i\phi}, Y_{\ell-m}(\theta, \phi) = (-1)^m Y_{\ell m}^*(\theta, \phi),$$

$$|N\rangle = |n\rangle - \sum_{m \neq n} |m\rangle \frac{1}{\epsilon_m - \epsilon_n} \langle m|V|n\rangle + \dots$$

$$E_N = \epsilon_n + \langle n|V|n\rangle - \sum_{m \neq n} \frac{|\langle m|V|n\rangle|^2}{\epsilon_m - \epsilon_n}$$

$$j_0(x) = \frac{\sin x}{x}, \quad n_0(x) = -\frac{\cos x}{x}, \quad j_1(x) = \frac{\sin x}{x^2} - \frac{\cos x}{x}, \quad n_1(x) = -\frac{\cos x}{x^2} - \frac{\sin x}{x}$$

$$j_2(x) = \left(\frac{3}{x^3} - \frac{1}{x}\right) \sin x - \frac{3}{x^2} \cos x, \quad n_2(x) = -\left(\frac{3}{x^3} - \frac{1}{x}\right) \cos x - \frac{3}{x^2} \sin x,$$

$$\frac{d}{dt} P_{i \rightarrow n}(t) = \frac{2\pi}{\hbar} |V_{ni}|^2 \delta(E_n - E_i),$$

$$\frac{d\sigma}{d\Omega} = \frac{m^2}{4\pi^2 \hbar^4} \left| \int d^3 r \mathcal{V}(r) e^{i(\vec{k}_f - \vec{k}_i) \cdot \vec{r}} \right|^2,$$

$$\sigma = \frac{(2S_R + 1) 4\pi (\hbar\Gamma_R/2)^2}{(2S_1 + 1)(2S_2 + 1) k^2 (\epsilon_k - \epsilon_r)^2 + (\hbar\Gamma_R/2)^2},$$

$$\frac{d\sigma}{d\Omega} = \left(\frac{d\sigma}{d\Omega}\right)_{\text{single}} \tilde{S}(\vec{q}), \quad \tilde{S}(\vec{q}) = \frac{1}{N} \left| \sum_{\vec{a}} e^{i\vec{q} \cdot \vec{a}} \right|^2 = \sum_{\delta\vec{a}} e^{i\vec{q} \cdot \delta\vec{a}},$$

$$\left(\frac{d\sigma}{d\Omega}\right)_{\text{Rutherford}} = \frac{e^4 Z_1^2 Z_2^2 m^2}{(\hbar k)^4 (1 - \cos \theta)^2}$$

$$\frac{d\sigma}{d\Omega} = \left(\frac{d\sigma}{d\Omega}\right)_{\text{Rutherford}} \left| \frac{1}{e} \int d^3 r \rho(\vec{r}) e^{i\vec{q} \cdot \vec{r}} \right|^2$$

$$e^{i\vec{k} \cdot \vec{r}} = \sum_{\ell} (2\ell + 1) i^{\ell} j_{\ell}(kr) P_{\ell}(\cos \theta),$$

$$P_{\ell}(\cos \theta) = \sqrt{\frac{4\pi}{2\ell + 1}} Y_{\ell, m=0}(\theta, \phi),$$

$$P_0(x) = 1, \quad P_1(x) = x, \quad P_2(x) = (3x^2 - 1)/3,$$

$$f(\Omega) \equiv \sum_{\ell} (2\ell + 1) e^{i\delta_{\ell}} \sin \delta_{\ell} \frac{1}{k} P_{\ell}(\cos \theta)$$

$$\psi_{\vec{k}}(\vec{r})|_{R \rightarrow \infty} = e^{i\vec{k} \cdot \vec{r}} + \frac{e^{ikr}}{r} f(\Omega),$$

$$\frac{d\sigma}{d\Omega} = |f(\Omega)|^2, \quad \sigma = \frac{4\pi}{k^2} \sum_{\ell} (2\ell + 1) \sin^2 \delta_{\ell}, \quad \delta \approx -ak$$

$$L_{\pm} |\ell, m\rangle = \sqrt{\ell(\ell + 1) - m(m \pm 1)} |\ell, m \pm 1\rangle,$$

$$\langle \tilde{\beta}, J, M | T_q^k | \beta, \ell, m_{\ell} \rangle = \langle JM | k, q, \ell, m_{\ell} \rangle \frac{\langle \tilde{\beta}, J || T^{(k)} || \beta, \ell, J \rangle}{\sqrt{2J + 1}},$$

$$n = \frac{(2s + 1)}{(2\pi)^d} \int_{k < k_f} d^d k, \quad d \text{ dimensions},$$

$$\{\Psi_s(\vec{x}), \Psi_{s'}^{\dagger}(\vec{y})\} = \delta^3(\vec{x} - \vec{y}) \delta_{ss'},$$

$$\Psi_s^{\dagger}(\vec{r}) = \frac{1}{\sqrt{V}} \sum_{\vec{k}} e^{i\vec{k} \cdot \vec{r}} a_s^{\dagger}(\vec{k}), \quad \{\Psi_s(\vec{x}), a_{\alpha}^{\dagger}\} = \phi_{\alpha, s}(\vec{x}).$$

1. Type α , β and γ particles exist in a ONE-DIMENSIONAL world. The α particle has mass M_α and is described by the one-dimensional field operator (in the interaction representation) within a large length L ,

$$\begin{aligned}\Phi_\alpha(\mathbf{x}, t) &= \frac{1}{\sqrt{L}} \sum_k e^{-iE_k t/\hbar + i\mathbf{k}\mathbf{x}} \mathbf{a}_k, \\ \Phi_\alpha^\dagger(\mathbf{x}, t) &= \frac{1}{\sqrt{L}} \sum_k e^{iE_k t/\hbar - i\mathbf{k}\mathbf{x}} \mathbf{a}_k^\dagger, \\ E_k &= M_\alpha c^2 + \frac{\hbar^2 \mathbf{k}^2}{2M_\alpha}.\end{aligned}$$

The β and γ particles are massless and described by the operators,

$$\begin{aligned}\Psi_\beta(\mathbf{x}, t) &= \frac{1}{\sqrt{L}} \sum_q e^{-iE_q t/\hbar + i\mathbf{q}\mathbf{x}} \mathbf{b}_q, \\ \Psi_\beta^\dagger(\mathbf{x}, t) &= \frac{1}{\sqrt{L}} \sum_q e^{iE_q t/\hbar - i\mathbf{q}\mathbf{x}} \mathbf{b}_q^\dagger, \\ \Psi_\gamma(\mathbf{x}, t) &= \frac{1}{\sqrt{L}} \sum_q e^{-iE_q t/\hbar + i\mathbf{q}\mathbf{x}} \mathbf{c}_q, \\ \Psi_\gamma^\dagger(\mathbf{x}, t) &= \frac{1}{\sqrt{L}} \sum_q e^{iE_q t/\hbar - i\mathbf{q}\mathbf{x}} \mathbf{c}_q^\dagger, \\ E_q &= \hbar c q.\end{aligned}$$

The massive α particle can decay to a β and a γ particle via the interaction

$$\mathbf{H}_{\text{int}} = g \int d\mathbf{x} \left[\Phi_\alpha(\mathbf{x}, t) \Psi_\beta^\dagger(\mathbf{x}, t) \Psi_\gamma^\dagger(\mathbf{x}, t) + \Phi_\alpha^\dagger(\mathbf{x}, t) \Psi_\beta(\mathbf{x}, t) \Psi_\gamma(\mathbf{x}, t) \right],$$

where the coupling constant g is small. The creation and destruction operators obey the commutation rules $[\mathbf{a}_k, \mathbf{a}_{k'}^\dagger] = \delta_{kk'}$, $[\mathbf{b}_q, \mathbf{b}_{q'}^\dagger] = \delta_{qq'}$ and $[\mathbf{c}_q, \mathbf{c}_{q'}^\dagger] = \delta_{qq'}$.

- (5 pts) Evaluate the commutator $[\Phi_\alpha(\mathbf{x}, t), \Phi_\alpha^\dagger(\mathbf{x}', t)]$.
- (5 pts) What is the dimensionality of g ?
- (25 pts) Calculate the rate at which an α particle at rest decays into a β and a γ particle.

Solution:

a)

$$\begin{aligned}
[\Phi_\alpha(x, t), \Phi_\alpha^\dagger(x', t)] &= \frac{1}{L} \sum_{kk'} e^{ikx - ik'x' - i\omega_k t + i\omega_{k'} t} [a_k, a_{k'}^\dagger] \\
&= \frac{1}{L} \sum_k e^{ik(x-x')} \\
&= \frac{1}{2\pi} \int dk e^{ik(x-x')} \\
&= \delta(x - x').
\end{aligned}$$

b)

$$\begin{aligned}
[E] &= [g][L][1/L^{3/2}], \\
[g] &= [E][L]^{1/2}.
\end{aligned}$$

c) First calculate matrix element

$$\begin{aligned}
\langle \beta, q; \gamma, q' | H_{\text{int}} | \alpha, P = 0 \rangle &= \frac{g}{L^{3/2}} \int dx \langle 0 | b_q c_{q'} \left(\sum_{kk'K} e^{-i(k+k'-K)x} b_k^\dagger c_{k'}^\dagger a_K \right) a_{P=0}^\dagger | 0 \rangle \\
&= \frac{g}{L^{3/2}} \int dx e^{-i(q+q')x} \\
&= \frac{g}{L^{1/2}} \delta_{q, -q'}.
\end{aligned}$$

Now, apply Fermi's golden rule.

$$\begin{aligned}
\Gamma &= \frac{2\pi}{\hbar} \frac{g^2}{L} \sum_{qq'} \delta_{q, -q'} \delta(Mc^2 - E_q - E_{q'}) \\
&= \frac{2\pi}{\hbar} \frac{g^2}{L} \sum_q \delta(Mc^2 - 2E_q), \\
&= \frac{g^2}{\hbar} \int_{-\infty}^{\infty} dq \delta(Mc^2 - 2\hbar c q) \\
&= \frac{2g^2}{\hbar} \frac{1}{2\hbar c} \\
&= \frac{g^2}{\hbar^2 c}.
\end{aligned}$$

2A (20 pts) Imagine you had calculated the following matrix element,

$$\mathcal{M} = \langle \alpha, J_f = 1, M_f = 0 | x^2 + y^2 - 2z^2 | \beta, J_i = 3, M_i = 0 \rangle.$$

For the following matrix elements, first state whether they are zero, and if not, express them in terms of \mathcal{M} and Clebsch-Gordan coefficients. **You do NOT need to evaluate any Clebsch-Gordan coefficients in your answers.**

- $\langle \alpha, J_f = 1, M_f = 0 | x^2 + y^2 + z^2 | \beta, J_i = 3, M_i = 0 \rangle$

- $\langle \alpha, J_f = 1, M_f = 0 | z^2 | \beta, J_i = 3, M_i = 0 \rangle$

- $\langle \alpha, J_f = 1, M_f = 1 | xz | \beta, J_i = 3, M_i = 1 \rangle$

- $\langle \alpha, J_f = 1, M_f = 1 | xy | \beta, J_i = 3, M_i = 2 \rangle$

Solution:

$$0 = \langle \alpha, J_f = 1, M_f = 0 | x^2 + y^2 + z^2 | \beta, J_i = 3, M_i = 0 \rangle$$

because the operator is a scalar.

$$\begin{aligned} 0 &= (x^2 + y^2 + z^2) = (x^2 + y^2 - 2z^2 + 3z^2) \\ z^2 &= (x^2 + y^2 + z^2)/3 - (x^2 + y^2 - 2z^2)/3, \\ \langle \alpha, J_f = 1, M_f = 0 | z^2 | \beta, J_i = 3, M_i = 0 \rangle \\ &= -\langle \alpha, J_f = 1, M_f = 0 | x^2 + y^2 - 2z^2 | \beta, J_i = 3, M_i = 0 \rangle / 3 \\ &= -\mathcal{M}/3. \end{aligned}$$

The last two matrix elements are zero.

$$\langle \alpha, J_f = 1, M_f = 1 | xz | \beta, J_i = 3, M_i = 1 \rangle = 0$$

because xz is comprised of $T_{\pm 1}^2$

$$\langle \alpha, J_f = 1, M_f = 1 | xy | \beta, J_i = 3, M_i = 2 \rangle$$

because xy is comprised of $T_{\pm 2}^2$

2B (15 pts) Now, imagine you had calculated a matrix element,

$$\mathcal{V} = \langle \alpha, J_f = 1, M_f = 0 | P_z | \beta, J_i = 0, M_i = 0 \rangle.$$

For all three values of M_f , express the following matrix elements in terms of \mathcal{V} and Clebsch-Gordan coefficients. **Again, You do NOT need to evaluate any Clebsch-Gordan coefficients in your answers.**

$\langle \alpha, J_f = 1, M_f | P_x | \beta, J_i = 0, M_i = 0 \rangle$. Be sure to note when a matrix element vanishes.

Solution:

$$\begin{aligned} P_z &= |P| \sqrt{\frac{4\pi}{3}} Y_{\ell=1, m=0}, \\ \sin \theta e^{\pm i\phi} &= \sqrt{\frac{8\pi}{3}} \mp Y_{1, \pm 1}, \\ P_x &= |P| \sqrt{\frac{2\pi}{3}} (-Y_{\ell=1, m=1} + Y_{\ell=1, m=-1}). \end{aligned}$$

The three elements are now:

$$\begin{aligned} \langle J_f = 1, M_f = 1 | P_x | J_i = 0, M_i = 0 \rangle &= \frac{-\mathcal{V}}{\sqrt{2}} \frac{\langle 1, 1 | 1, 1, 0, 0 \rangle}{\langle 1, 0 | 1, 0, 0, 0 \rangle}, \\ \langle J_f = 1, M_f = -1 | P_x | J_i = 0, M_i = 0 \rangle &= \frac{\mathcal{V}}{\sqrt{2}} \frac{\langle 1, -1 | 1, -1, 0, 0 \rangle}{\langle 1, 0 | 1, 0, 0, 0 \rangle}, \\ \langle J_f = 1, M_f = 0 | P_x | J_i = 0, M_i = 0 \rangle &= 0. \end{aligned}$$

3 In the center of a ONE-DIMENSIONAL star, matter consists of weakly interacting electrons, protons and neutrons. The baryon density is n_B baryons per length. Interactions of the type

$$n \rightarrow e + p + \bar{\nu}, \quad e + p \rightarrow n + \nu,$$

can proceed with the neutrinos leaving the star until the energy is minimized for the given n_B . Assume the neutrons and protons have the same mass M , and can be treated non-relativistically, whereas the electrons can be treated as if they are massless. Refer to the three Fermi wave numbers as k_e , k_n and k_p .

- (a) (15 pts) Write three equations involving k_e , k_n and k_p , which when solved can yield the three Fermi wave numbers.
- (b) (15 pts) Solve for the neutron fraction, i.e. the fraction of baryons that are neutrons.

Solution:

a) 1. Charge conservation: $k_e = k_p$

One-dimensional density,

$$n = \frac{2}{\pi} \int_0^{k_f} dk = \frac{2}{\pi} k_f.$$

2. Baryon conservation $n_B = \frac{2}{\pi}(k_n + k_p)$

3 Minimize energy

$$\frac{\hbar^2 k_n^2}{2M} = \frac{\hbar^2 k_p^2}{2M} + \hbar c k_e.$$

b)

$$\begin{aligned} \frac{\hbar^2}{2M} \left[\left(\frac{\pi n_B}{2} - k_p \right)^2 - k_p^2 \right] - \hbar c k_p &= 0, \\ \frac{\hbar^2}{2M} \left[-\pi n_B k_p + \left(\frac{\pi n_B}{2} \right)^2 \right] - \hbar c k_p &= 0, \\ k_p \left(\frac{\hbar^2}{2M} \pi n_B + \hbar c \right) &= \frac{\hbar^2}{2M} \left(\frac{\pi n_B}{2} \right)^2, \end{aligned}$$

$$\begin{aligned}
k_p &= \frac{\frac{\hbar^2}{2M} \left(\frac{\pi n_B}{2}\right)^2}{\left(\frac{\hbar^2}{2M} \pi n_B + \hbar c\right)} \\
&= \frac{\pi n_B / 2}{2 + 4Mc / (\pi \hbar n_B)}, \\
k_n &= \frac{\pi}{2} n_B - k_p \\
&= \frac{\pi n_B}{2} \left(1 - \frac{1}{2 + 4Mc / (\pi \hbar n_B)}\right) \\
&= \frac{\pi n_B}{2} \frac{1 + 4Mc / (\pi \hbar n_B)}{2 + 4Mc / (\pi \hbar n_B)} \\
\frac{n_n}{n_B} &= \frac{1 + 4Mc / (\pi \hbar n_B)}{2 + 4Mc / (\pi \hbar n_B)}.
\end{aligned}$$